

The Canadian High Arctic: a Window into Climate, Ozone & Global Air Quality



Dan Weaver

Sigma Xi Distinguished Lecture Series

19 November 2015

Introductions



Dan Weaver

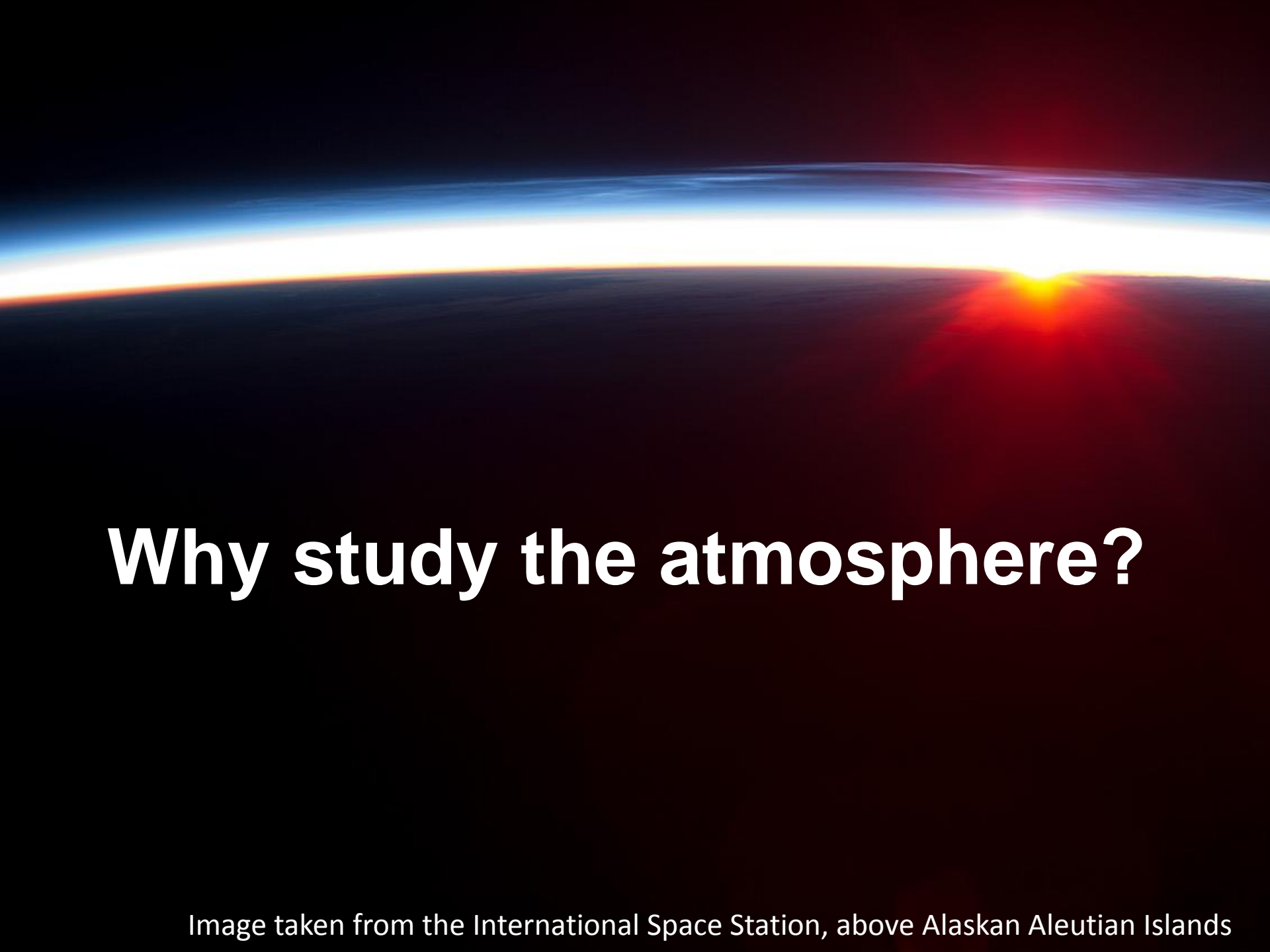


- Ph.D. Candidate, Dept. of Physics
- Travel to PEARL for annual winter/spring campaigns (since 2012)
- I also run campaign communications

Motivation for atmospheric research in the high Arctic?

Understanding planetary
atmospheric and environmental
change

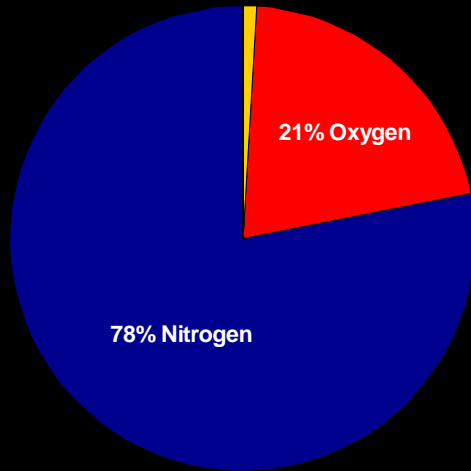




Why study the atmosphere?

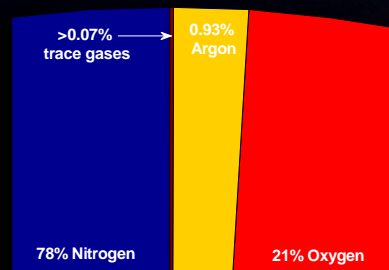
Image taken from the International Space Station, above Alaskan Aleutian Islands

What is air?



Trace Gases include:

- Argon
- Water vapour
- Carbon dioxide
- Methane
- Nitrous oxide
- Ozone
- Many others



Small amounts of gases can have a BIG impact on the planet!

Why is this research important
now?





1793

Painting by Elizabeth Simcoe,
from the City of Toronto "*The History of Toronto: An 11,000-Year Journey*" web page

Downtown Toronto

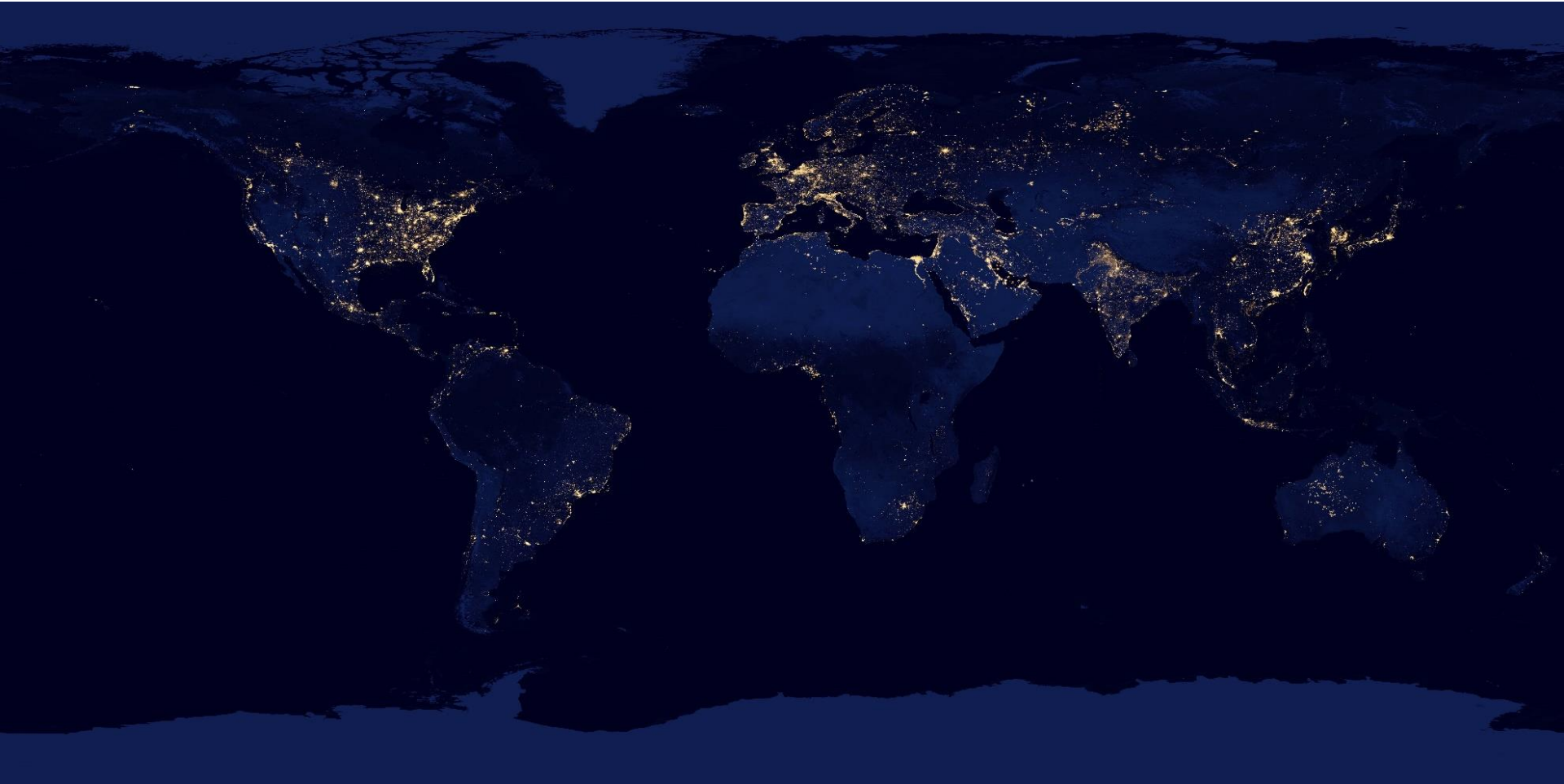


1890s



2014

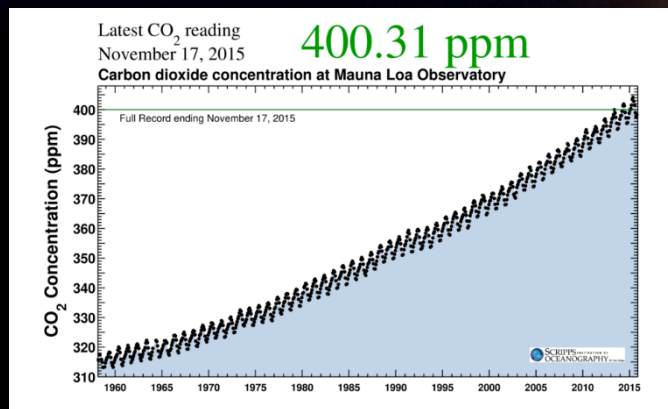
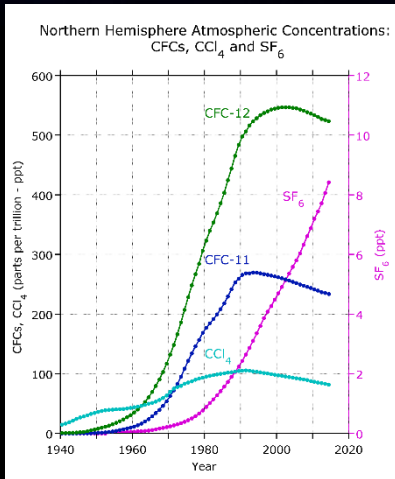
Earth at night, 2015



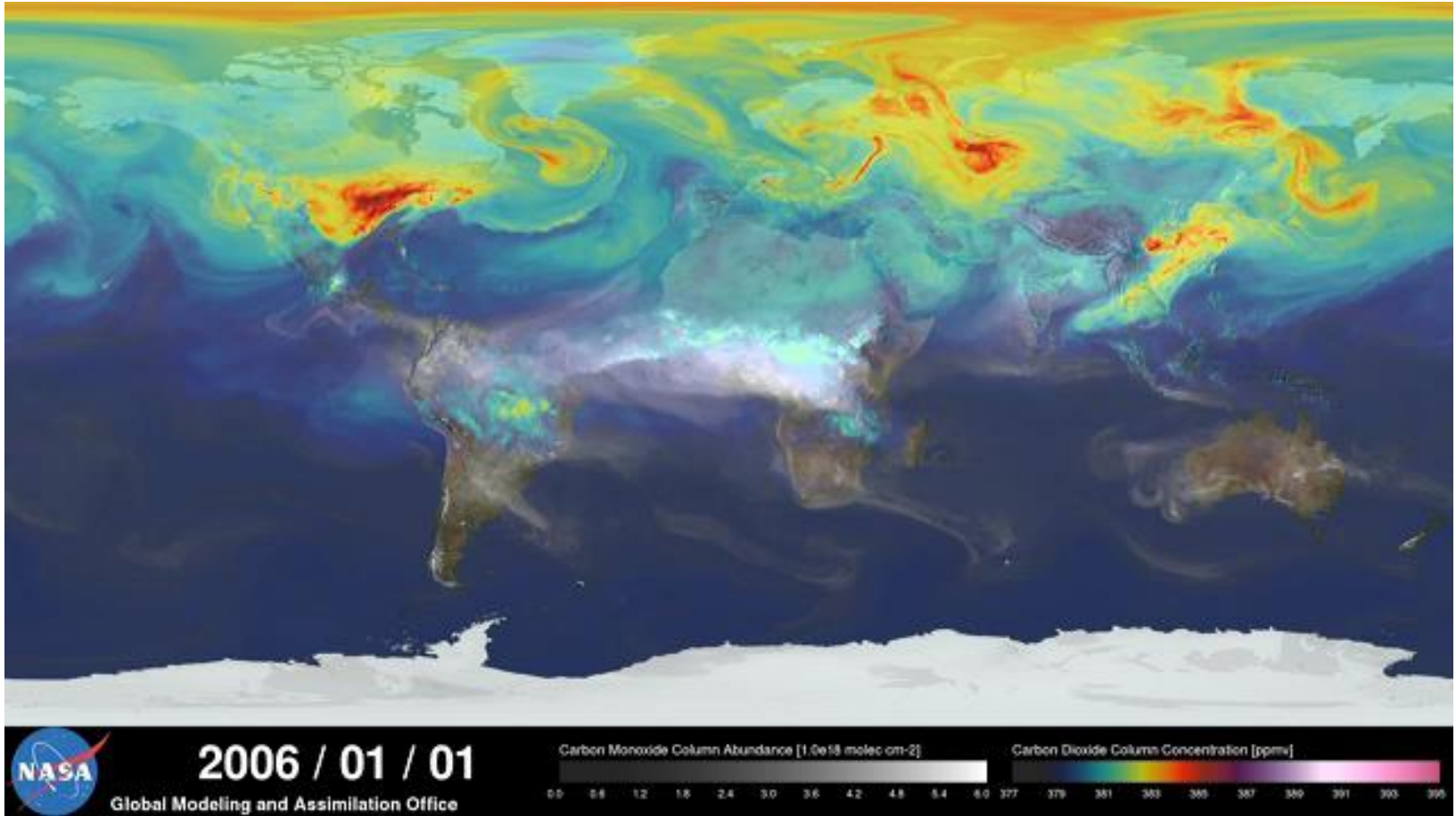
Credit: NASA

Atmospheric science

aims to measure & understand the changing atmosphere

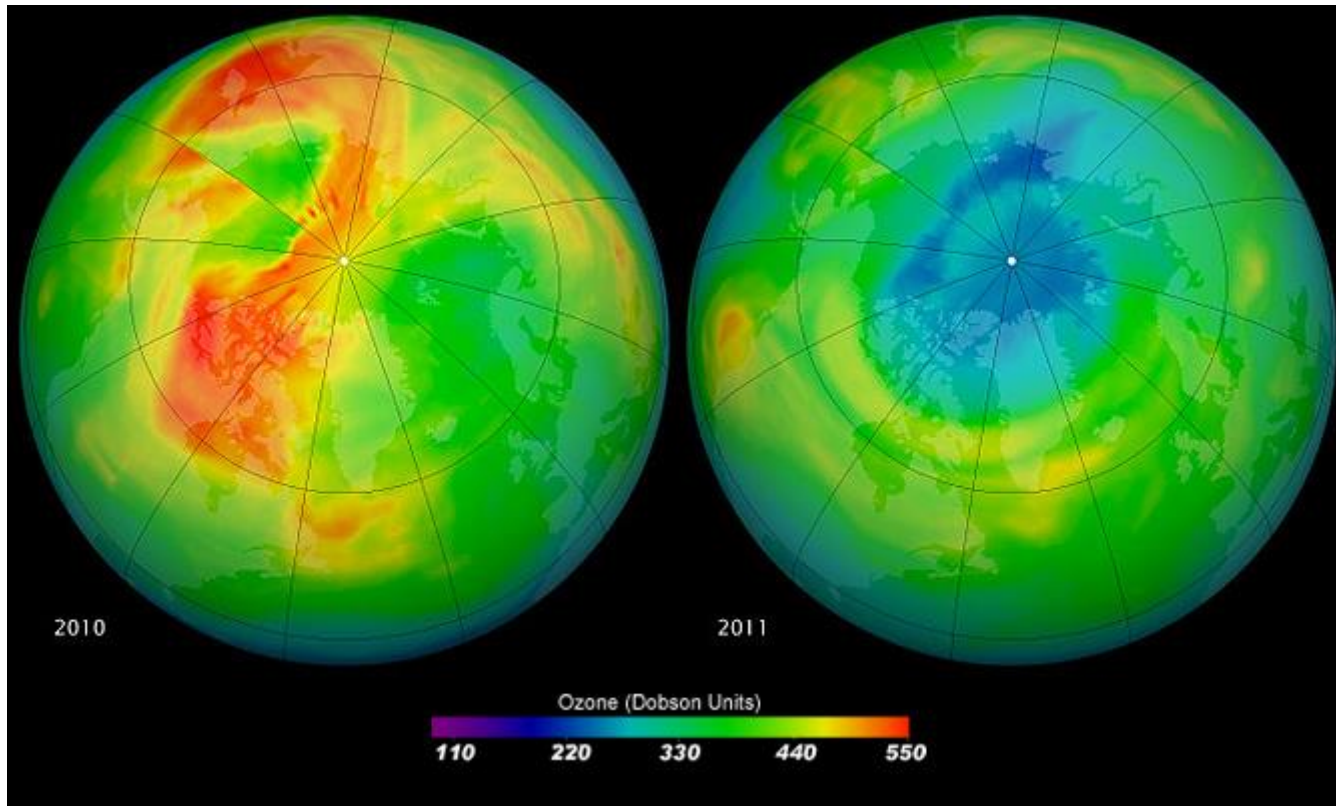


CO₂

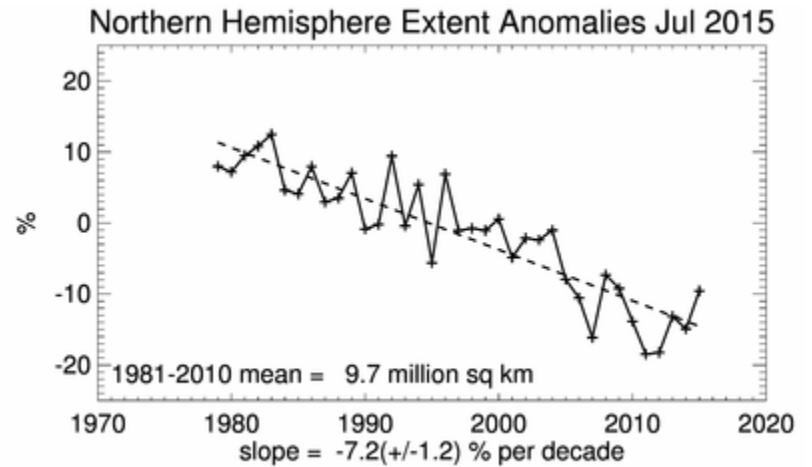


Arctic ozone change

Ozone depletion



Arctic sea ice change



Dwindling Arctic Sea Ice

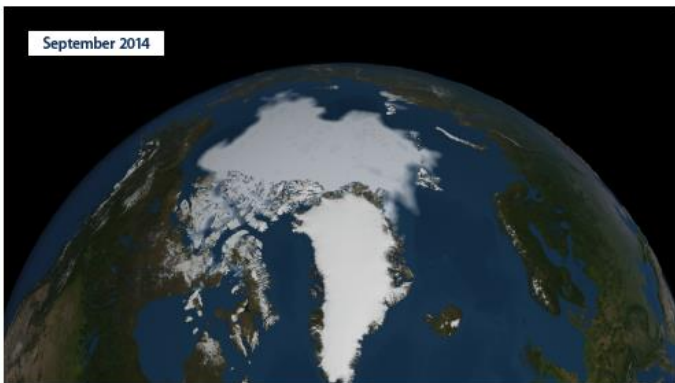
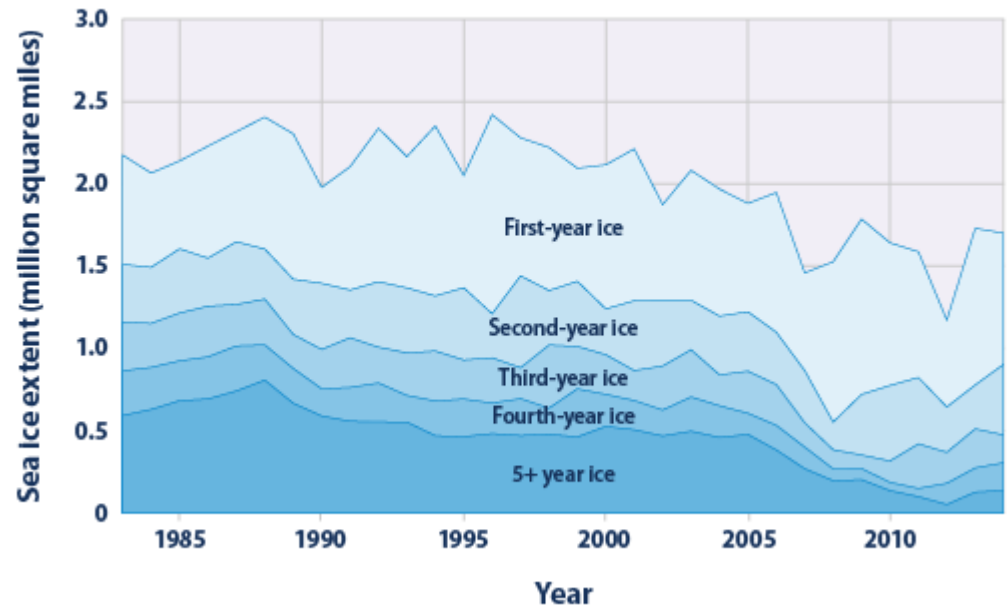


Figure 2. Age of Arctic Sea Ice at Minimum September Week, 1983–2014





**Our society's impressive rapid technological development
has come at a significant **environmental cost****

**we should at least
understand how we are changing the planet**

Ideally, we should avoid them entirely

A scenic photograph of a sunset or sunrise over a snowy mountain range. The sun is a bright, glowing orb in the upper right, casting a warm orange and yellow light across the sky. Below the sun, the sky transitions into a deep blue. The mountains are covered in snow, with some peaks catching the low light of the sun. In the foreground, the dark, silhouetted tops of evergreen trees are visible against the blue sky.

Today's talk

**Where is Eureka
PEARL research
Working in the High
Arctic**



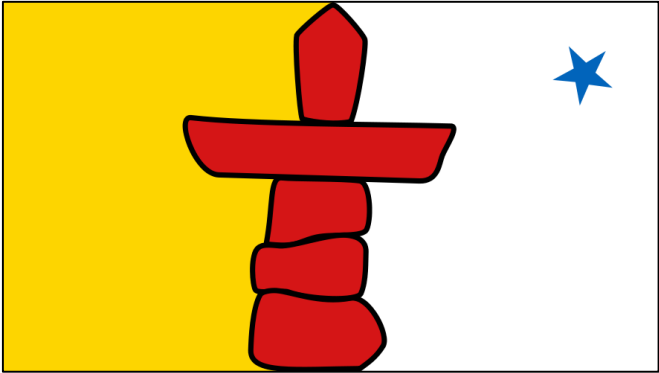
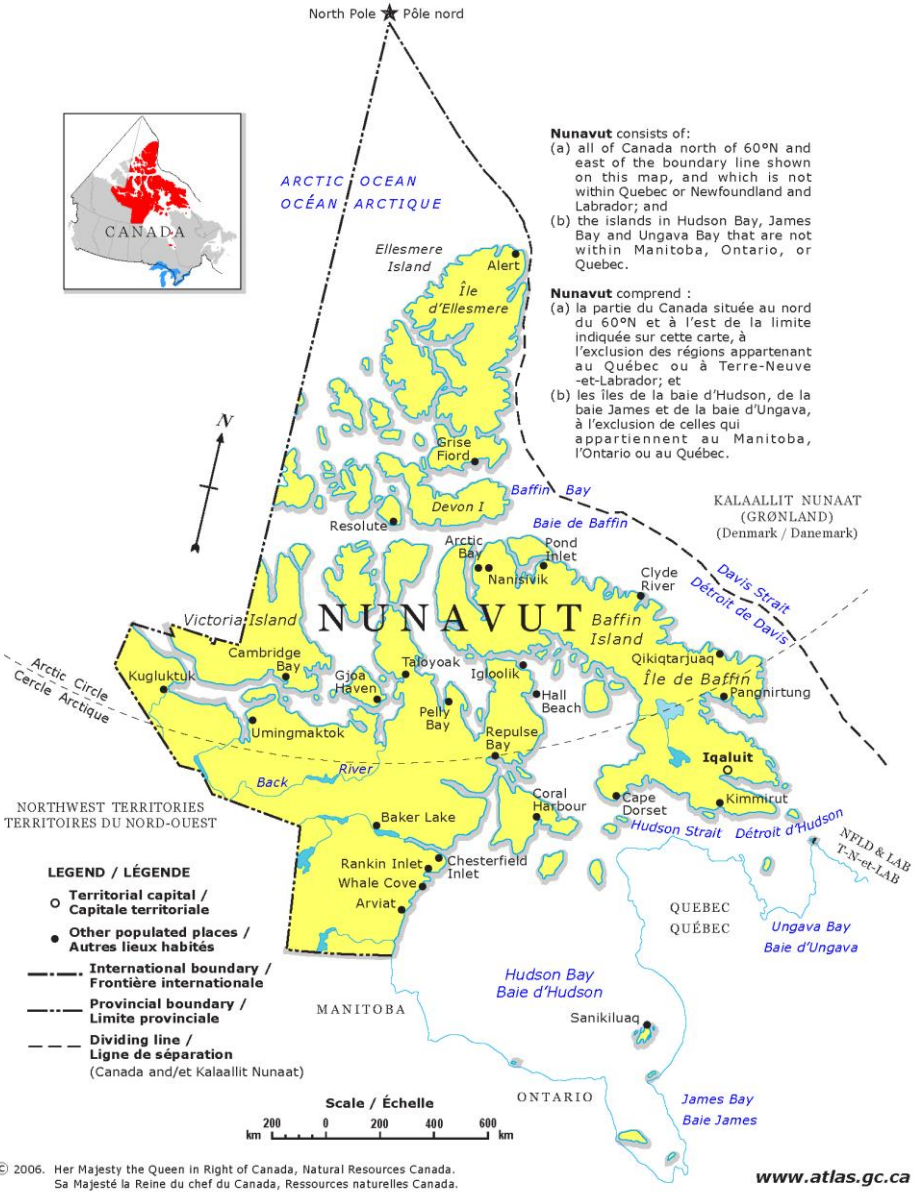
Canadian Arctic

Our Arctic is 40% of Canada's landmass & is home to over 100,000 Canadians.



Nunavut

- Territory created in 1999
- Over 80% Inuit population (total: ~37,000)
- Smallest economy, largest area of all provinces/territories
- Elects a single MP
- Has distinct government system (consensus)
- Unique culture (Inuit)





Eureka & PEARL

Eureka, Nunavut is at 80°N, 86.4°W on Ellesmere Island

Resolute

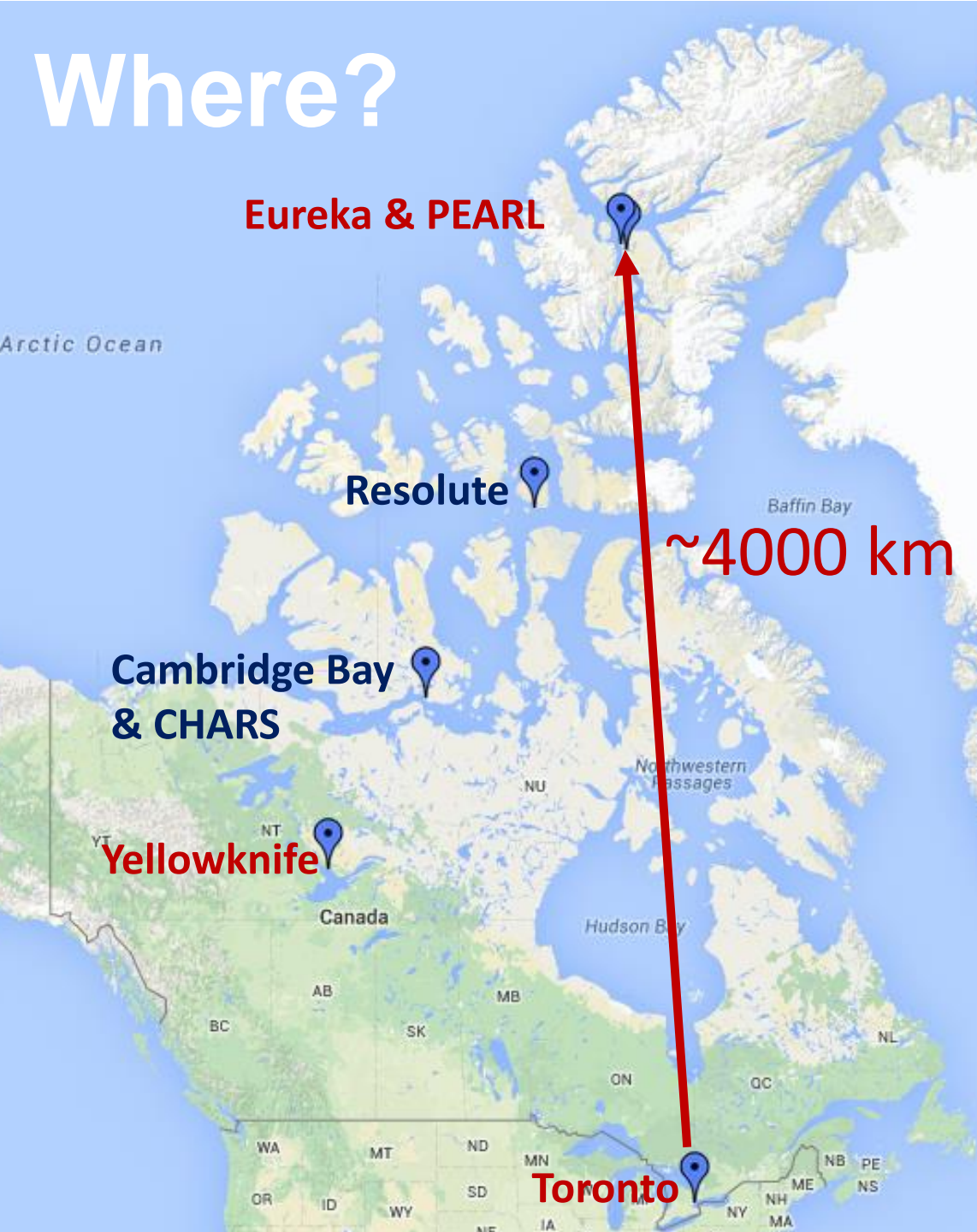
Cambridge Bay & CHARS

Yellowknife

Toronto

How far North?!

Where?



Eureka & PEARL

Resolute

**Cambridge Bay
& CHARS**

Yellowknife

Toronto

~4000 km

Eureka, Nunavut is
at 80°N
on Ellesmere Island

Distances

Eureka/PEARL is over
4000 km north of
Toronto!

*Yellowknife is ~halfway
point*



**Eureka is
closer to:**

Stockholm, Sweden
St. Petersburg, Russia
Reykjavik, Iceland
Glasgow, Scotland



**Than to
Toronto!**

Eureka is farther north than most Northern Lights!



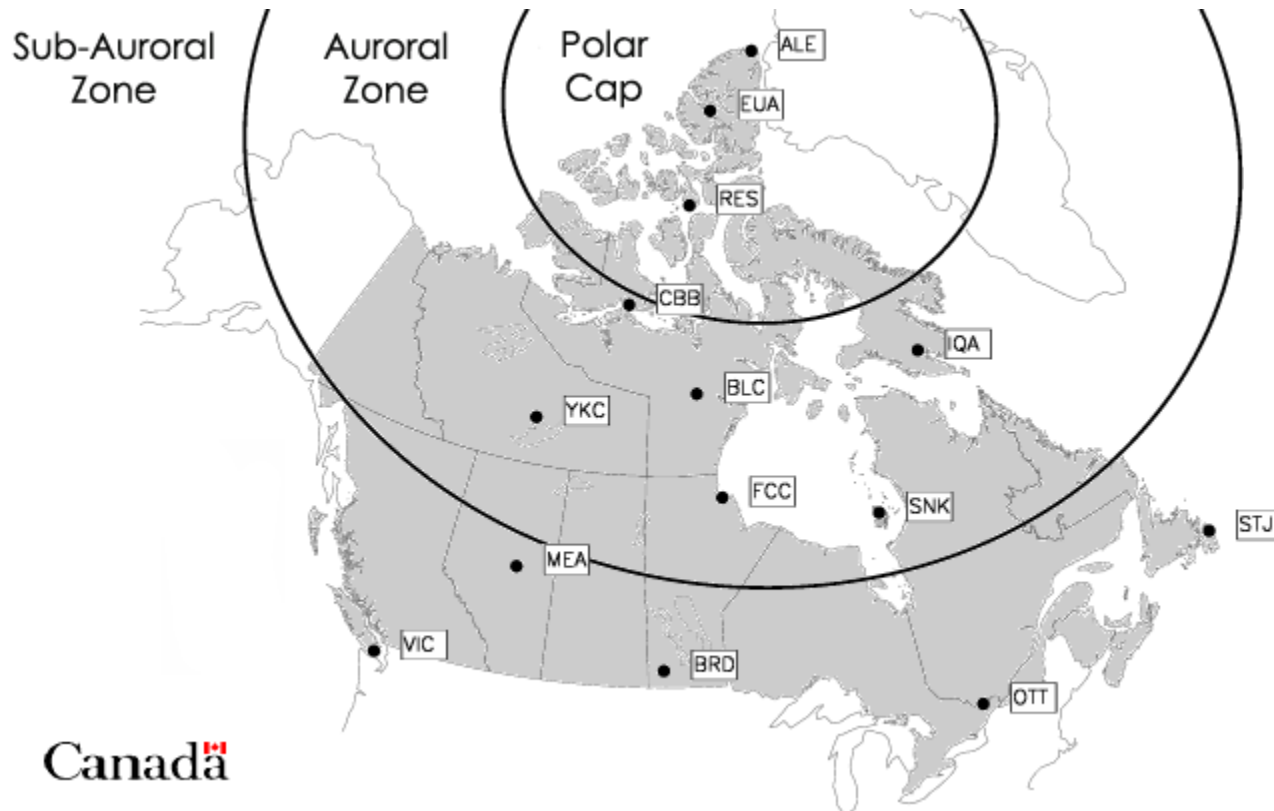
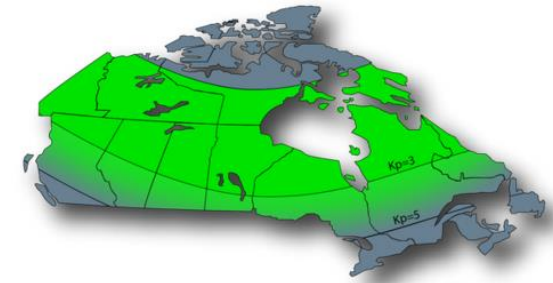
AuroraMAX @AuroraMAX - Nov 17

AURORAMAX ALERT • Observers in YT NT NU AB SK MB and northern BC ON QC NL be on alert for #auroras Nov 18-19, 2015.

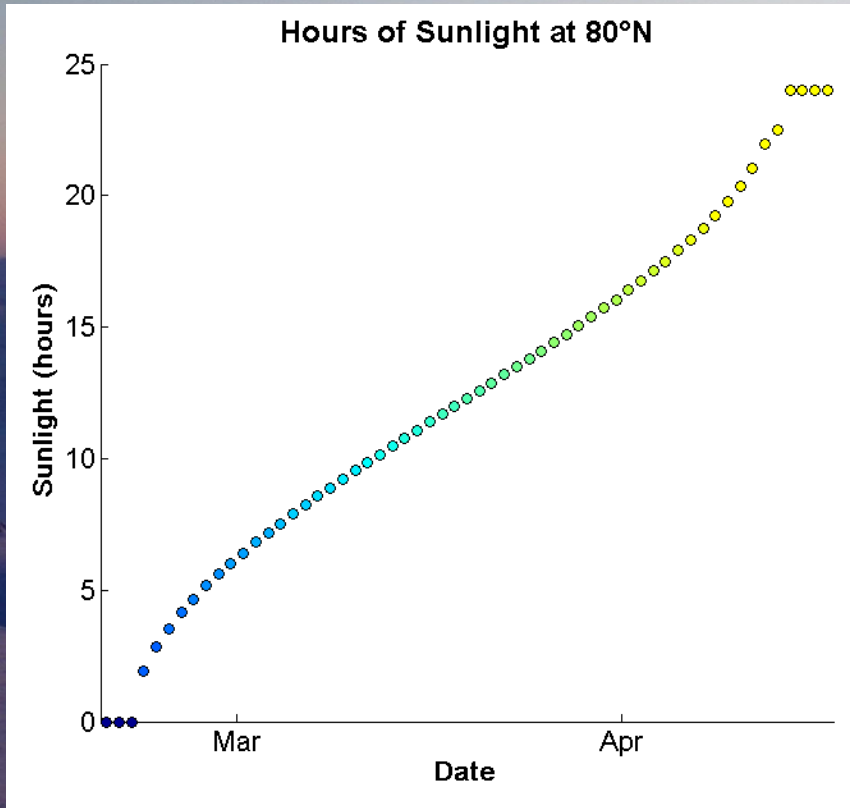


ALERT

BE ON ALERT FOR ACTIVE AURORAS!



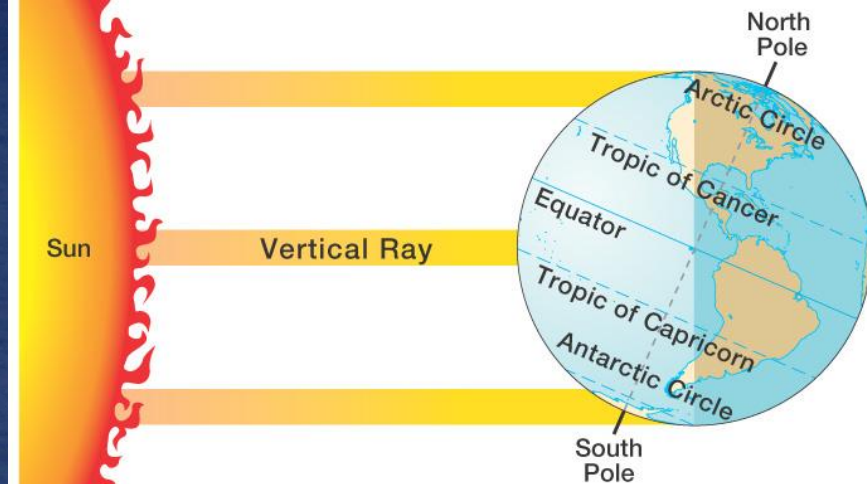
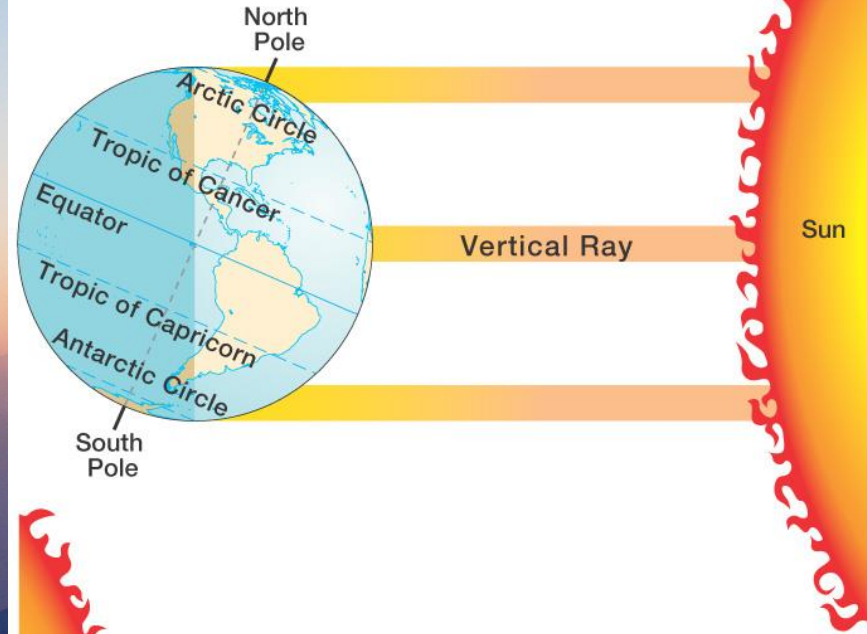
Daylight



Eureka, Nunavut
18 February – 18 April 2015
Data from the U.S. Naval Observatory

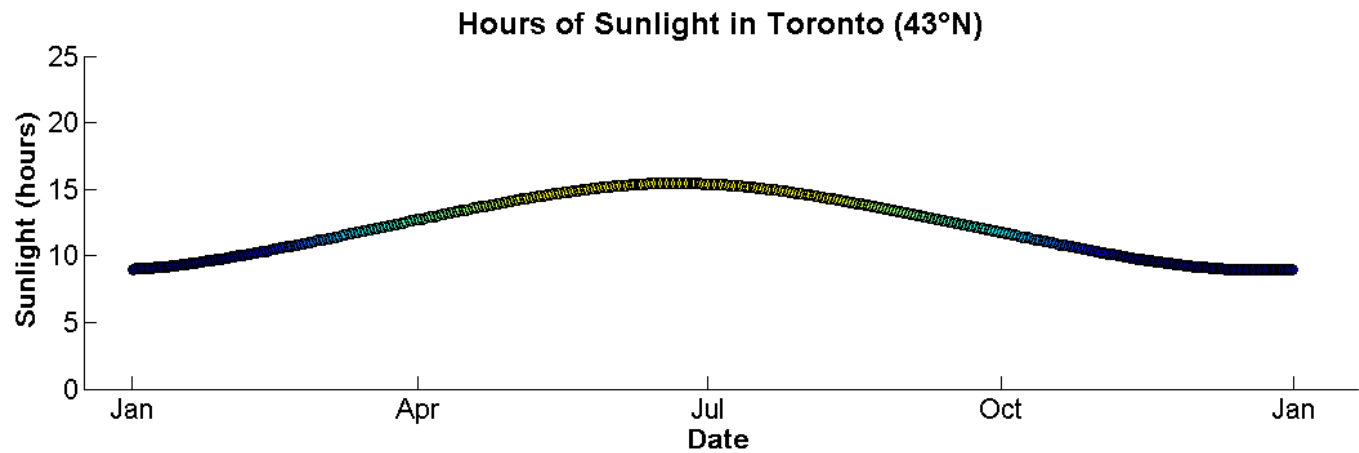
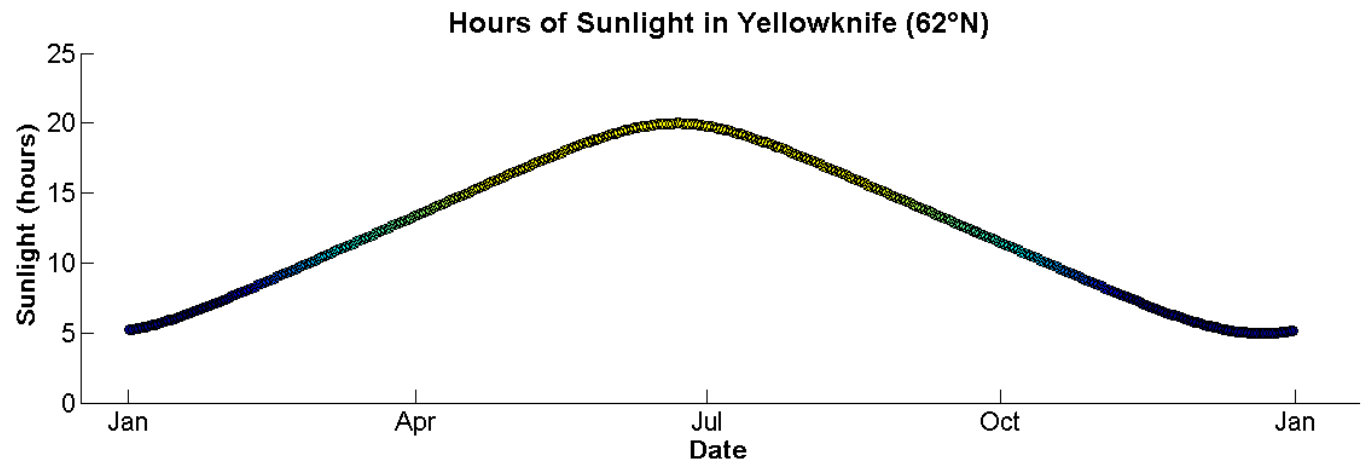
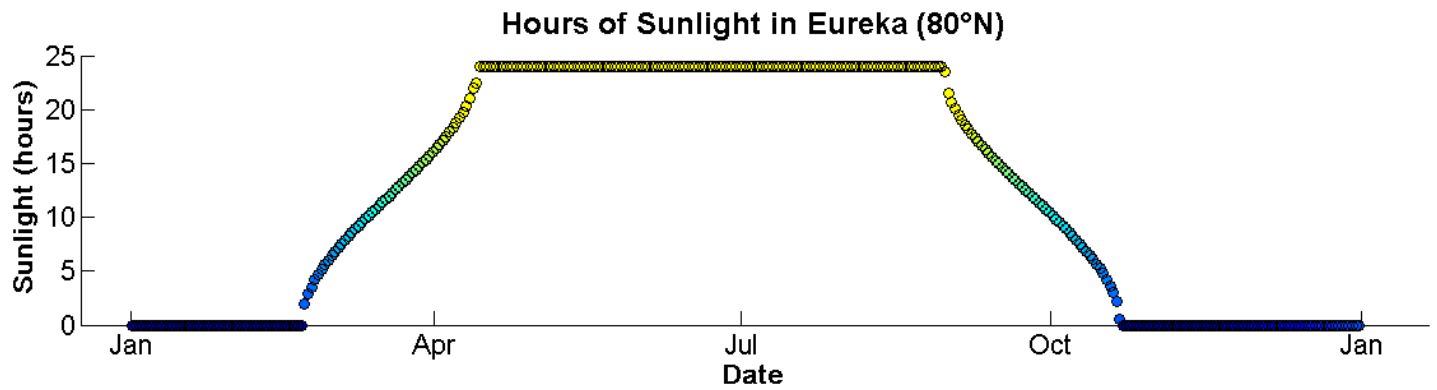
Side Views

June 21



December 21

Side View



Travelling to PEARL...

Toronto
Edmonton
Yellowknife
Cambridge Bay
Resolute
Eureka



Eureka International Airport





No Service 99% 11:13 PM 45%

Current **-47°C**
Clear Feels Like: -47

Wind Speed & Direction: **Calm** Humidity: —

Pressure: — Visibility: **16km** Ceiling: **unlimited**

Sunrise/Sunset: — Yesterday's High/Low: **-46.8° -49.3°**

Evening Sunday **-36°**
Weather delays unlikely

Overnight Sunday **-36°**
Weather delays unlikely

Morning Monday **-35°**
Weather delays unlikely

Monday **000 000**

The Weather Network **Eureka** NU, Canada

Dress warmly!



Eureka

- Environment Canada facility
- 8 permanent staff
- Short-term research teams



Eureka at twilight

Hydrogen shed



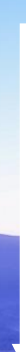
Eureka
Weather
Station



Diesel &
power
generation



Zero-altitude
PEARL
auxiliary lab
(OPAL)



The Eureka Weather Station



WELCOME TO EUREKA
WEATHER STATION

CANADA POSTES
POST CANADA
EUREKA, NU
XOA OGO



Eureka is the northernmost permanently settled civilian place in the world*



*Nord, Greenland and Alert, Nunavut are further north, but are primarily military

It's like a tiny town. There is:

- Post office
- Store
- Bar (BYOB)
- Airport
- Labs
- Guest rooms
- Power facilities
- Sewage facilities
- Water infrastructure
- ~~Hospital/clinic~~

Eureka Weather Station's recreation room



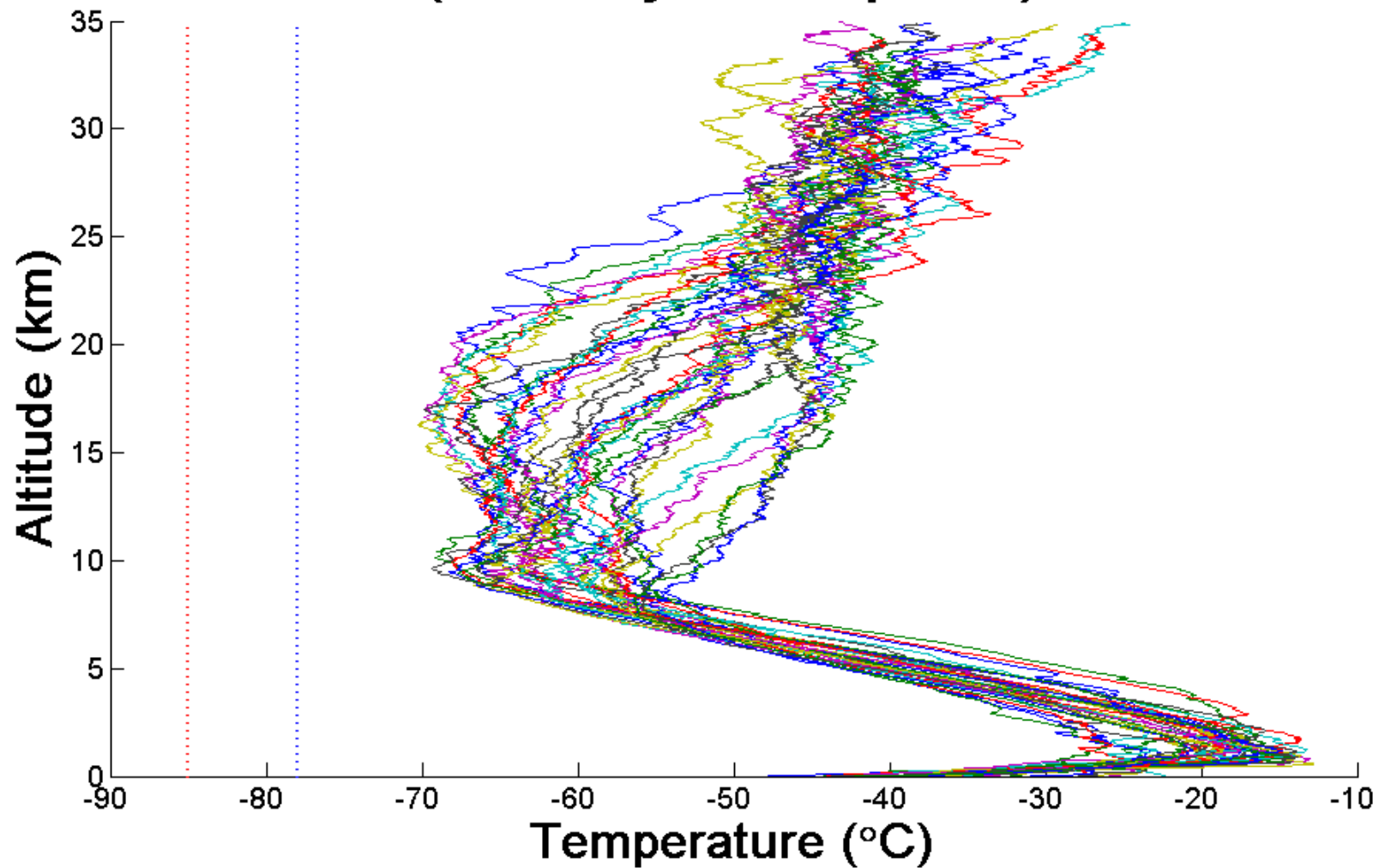
There's always plenty of food in Eureka

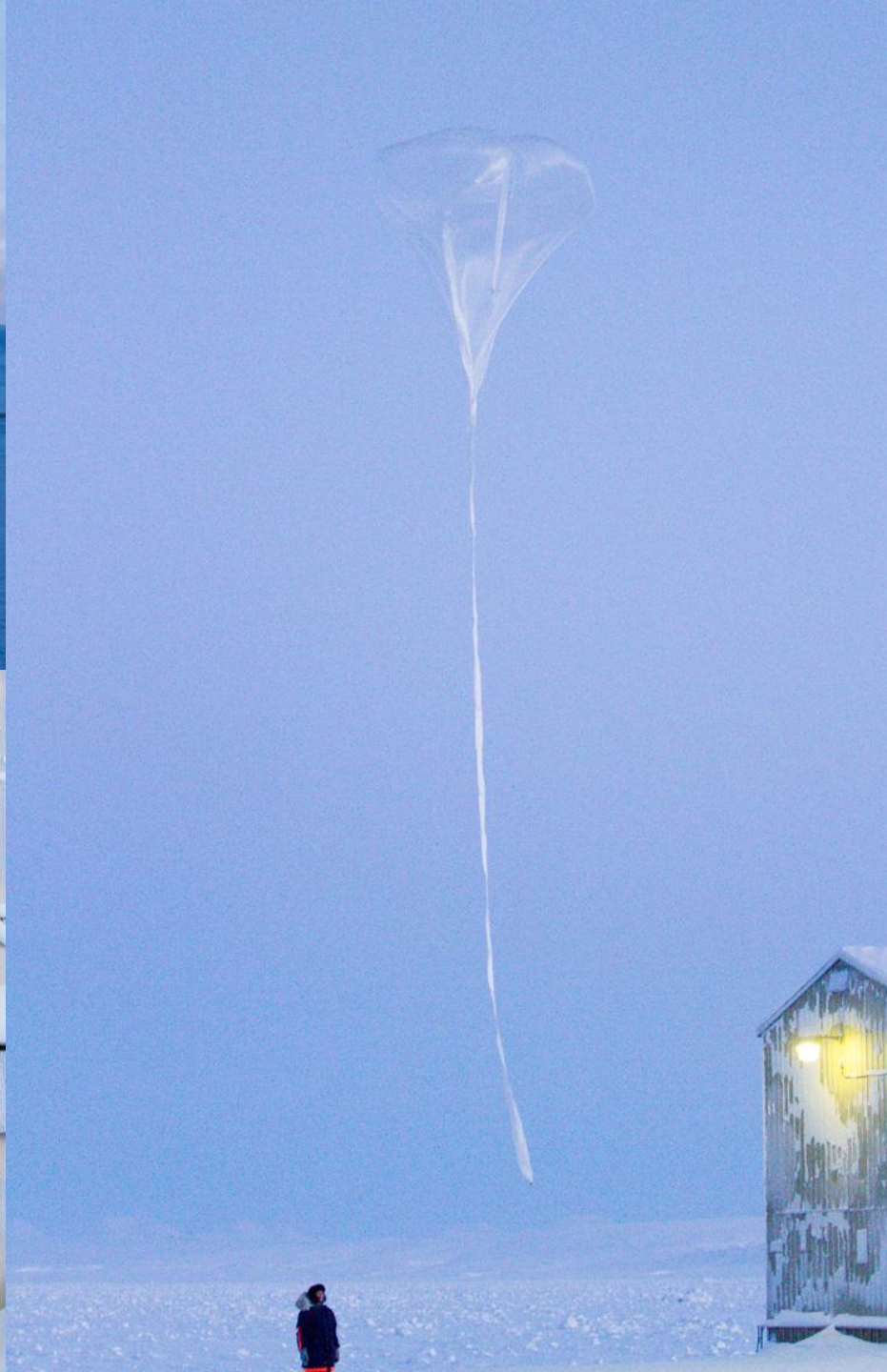


Diesel power generators & temperature inversions

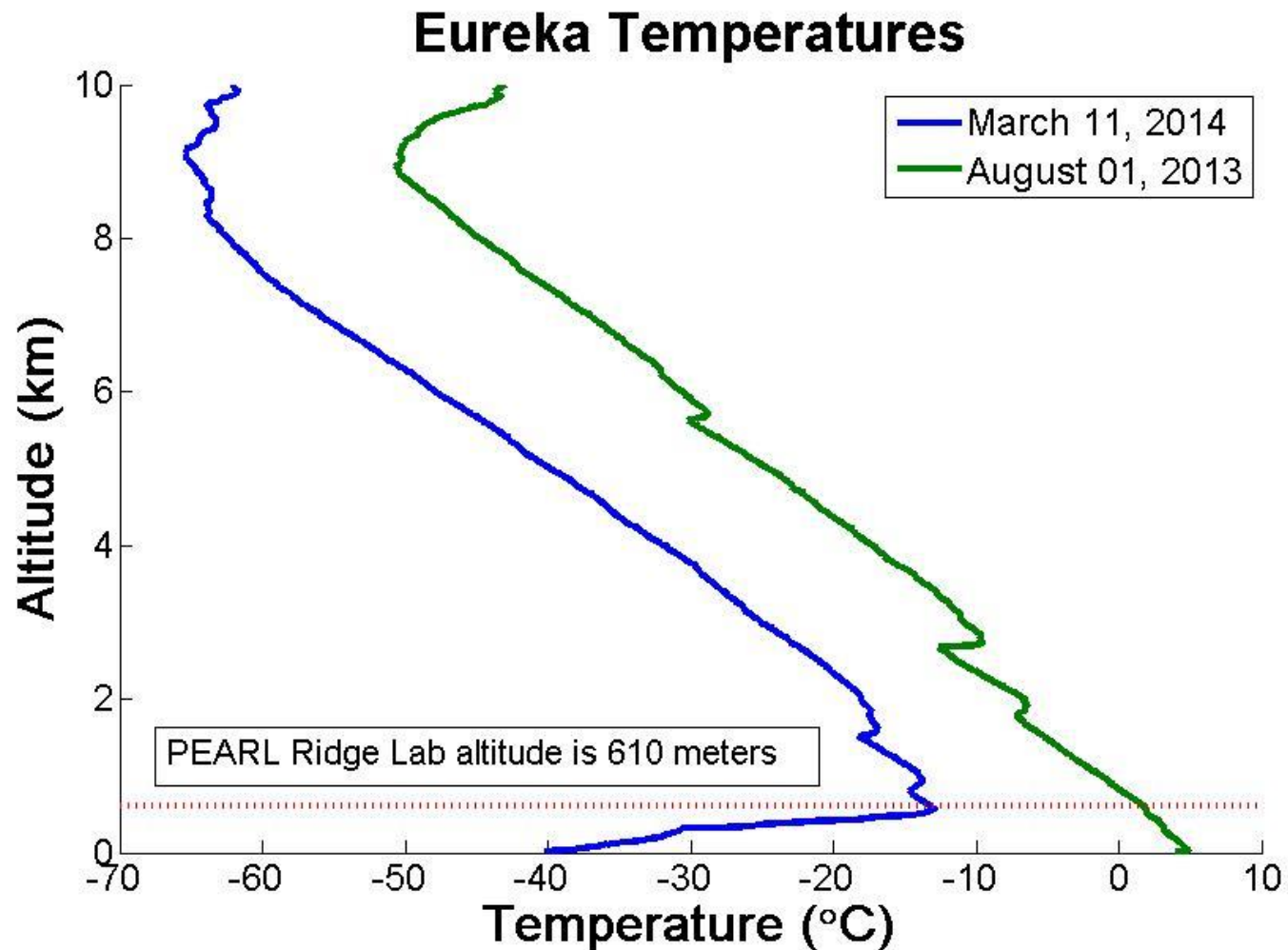


2014 ACE Campaign radiosonde profiles (February 26 to April 02)





Interesting phenomena: Temperature inversion











Driving to the PEARL Ridge Lab

Be prepared to shovel
snow when in the
Arctic!

DYK:
Canada's high Arctic is a
polar desert.
There is very low
humidity!



Figure .: Location of PEARL relative to Eureka.



EC has a variety of specialized vehicles...





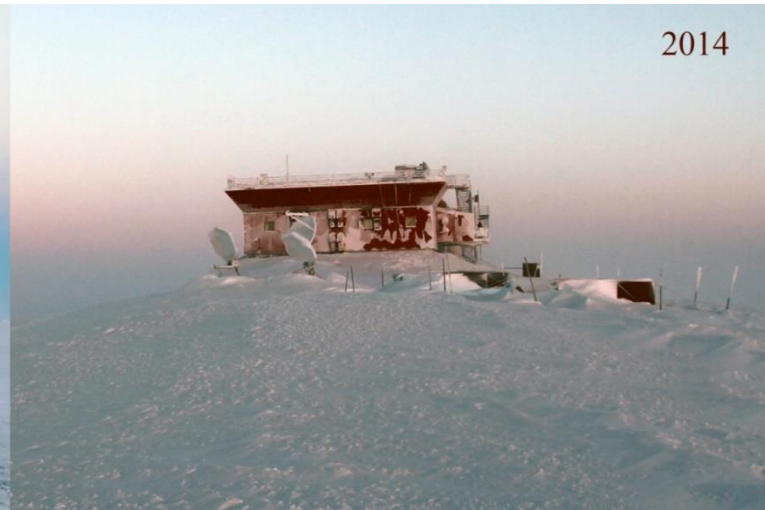
Along the road to the PEARL Ridge Lab



PEARL



- Ideally situated to study the high Arctic atmosphere
- Run by CANDAC since 2006
- Originally built as an Environment Canada facility, AStrO, in the 1990s (Arctic Stratospheric Ozone Observatory)



- 80.05° North, 86.4° West
- 610 meters a.s.l.
- ~15 km from Environment Canada's Eureka Weather Station

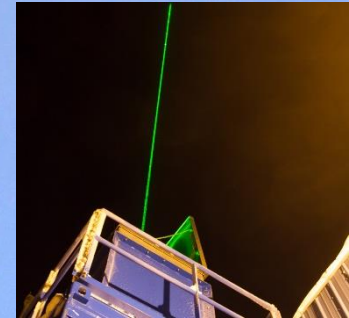
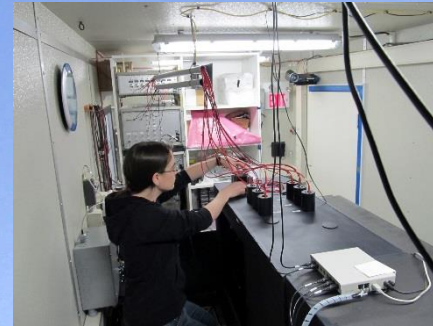
The PEARL Ridge Lab





It's not easy to get here, but it's beautiful!

Zero-altitude PEARL Auxiliary Lab (0PAL)



SAFIRE

(Surface and Atmospheric Flux, Irradiance and Radiation Extension)



The flux tower





Research at PEARL?

Ozone Depletion

Climate Change

Pollution

How do we measure the atmosphere?

Ground-based spectrometers



Ground-based lidars

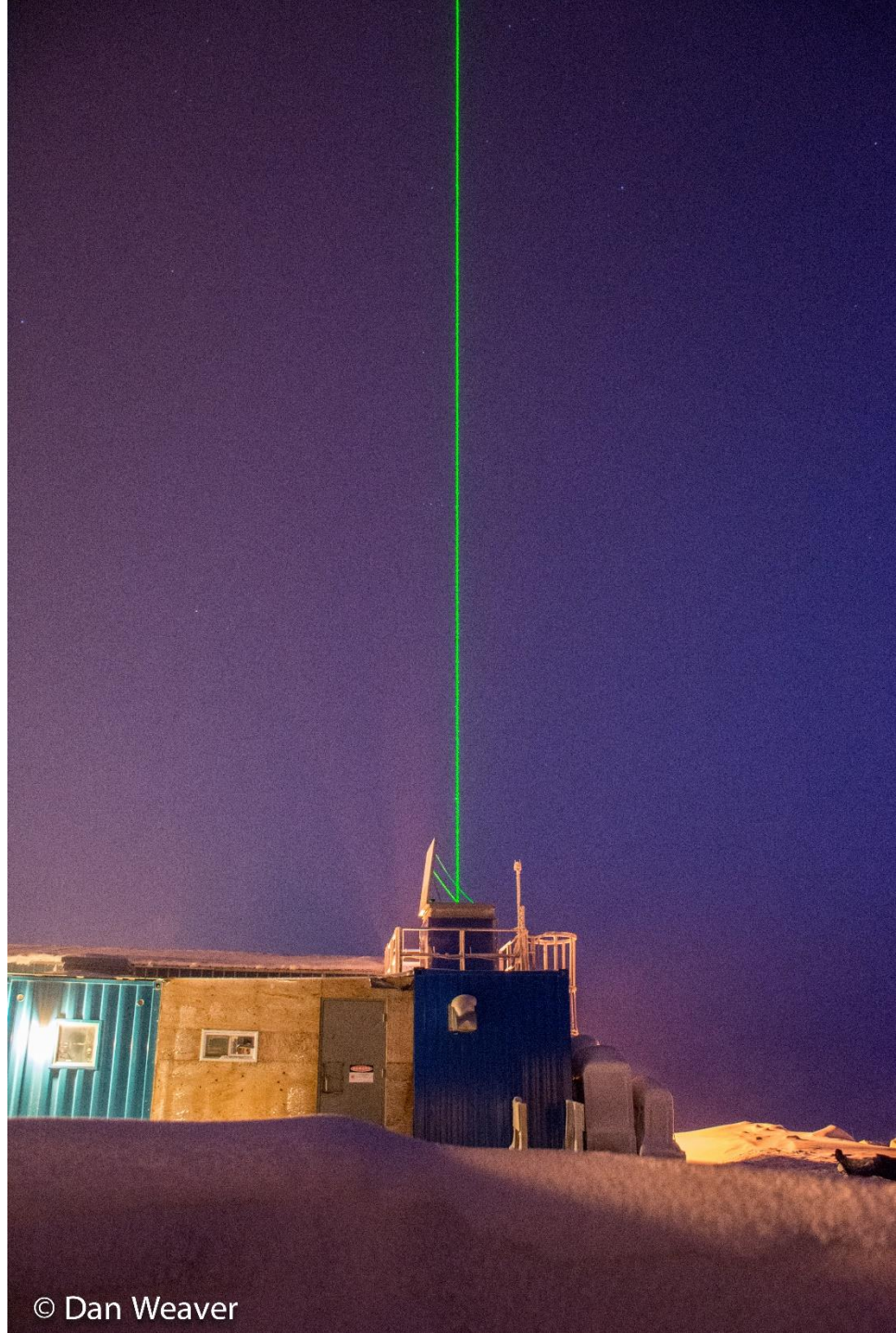
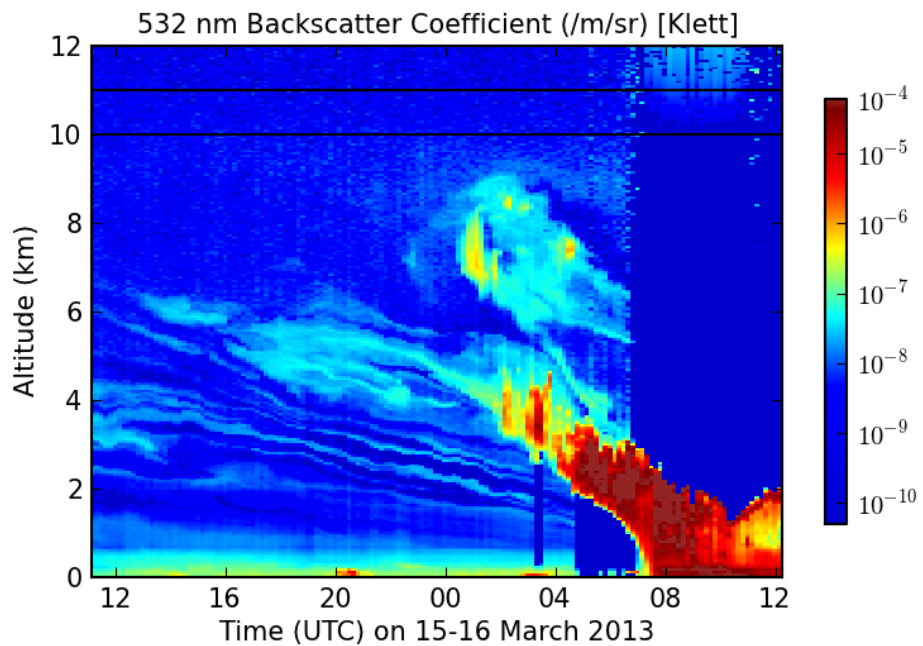


Balloon measurements

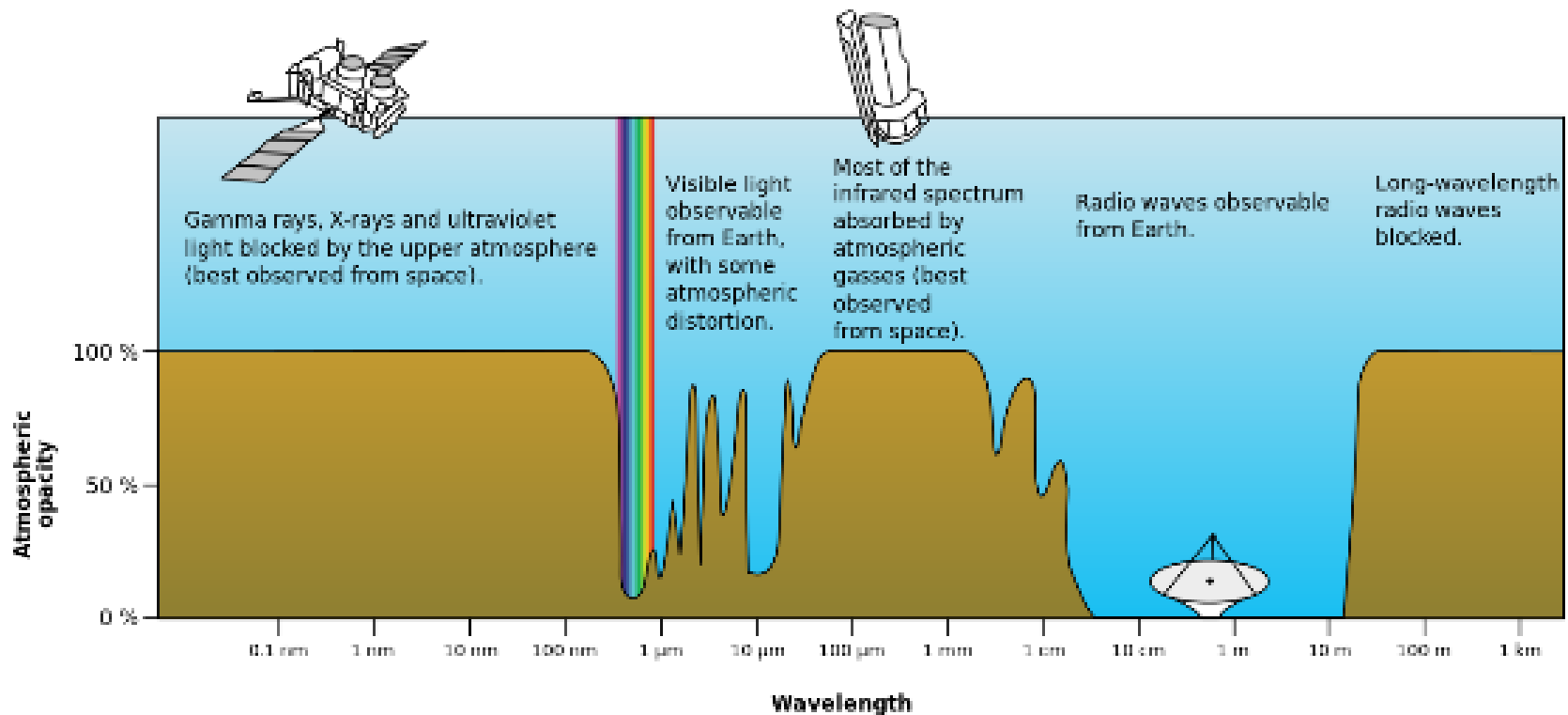


LIDARs

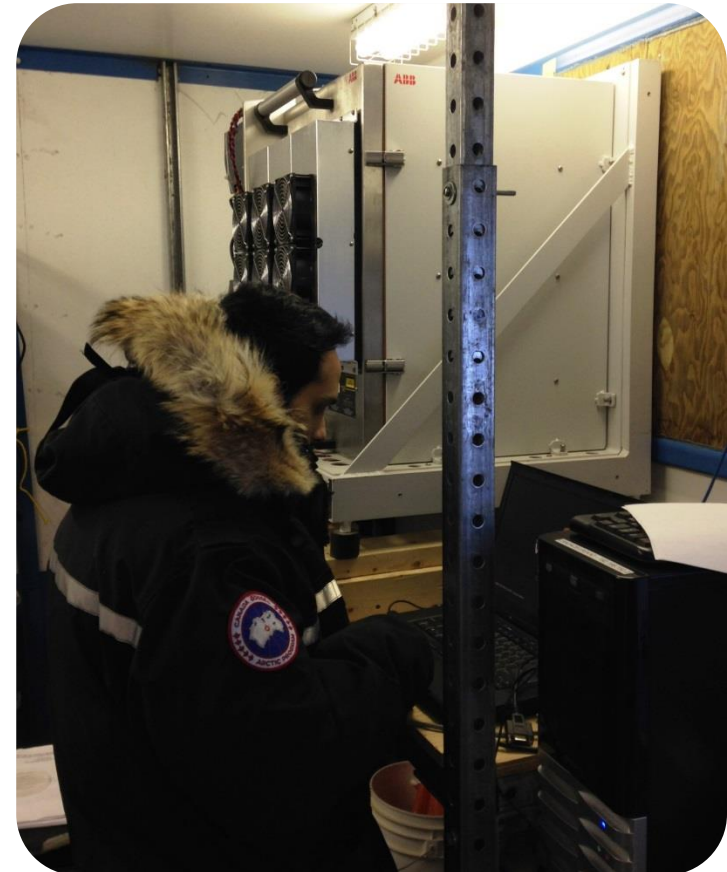
Light detection and ranging



PEARL's instruments mostly observe
the atmosphere
in the UV, visible, infrared and microwave regions
of the EM spectrum



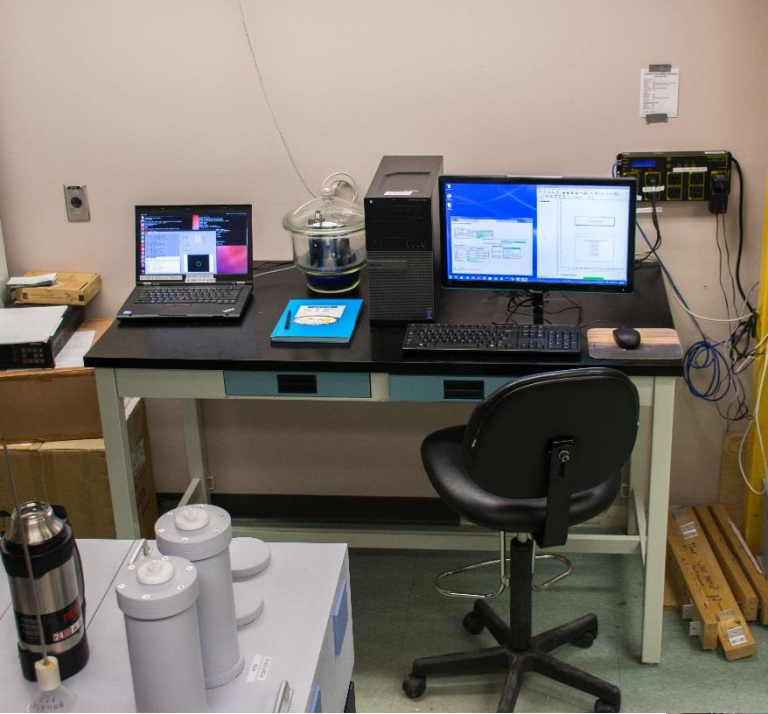
E-AERI: an emission FTIR



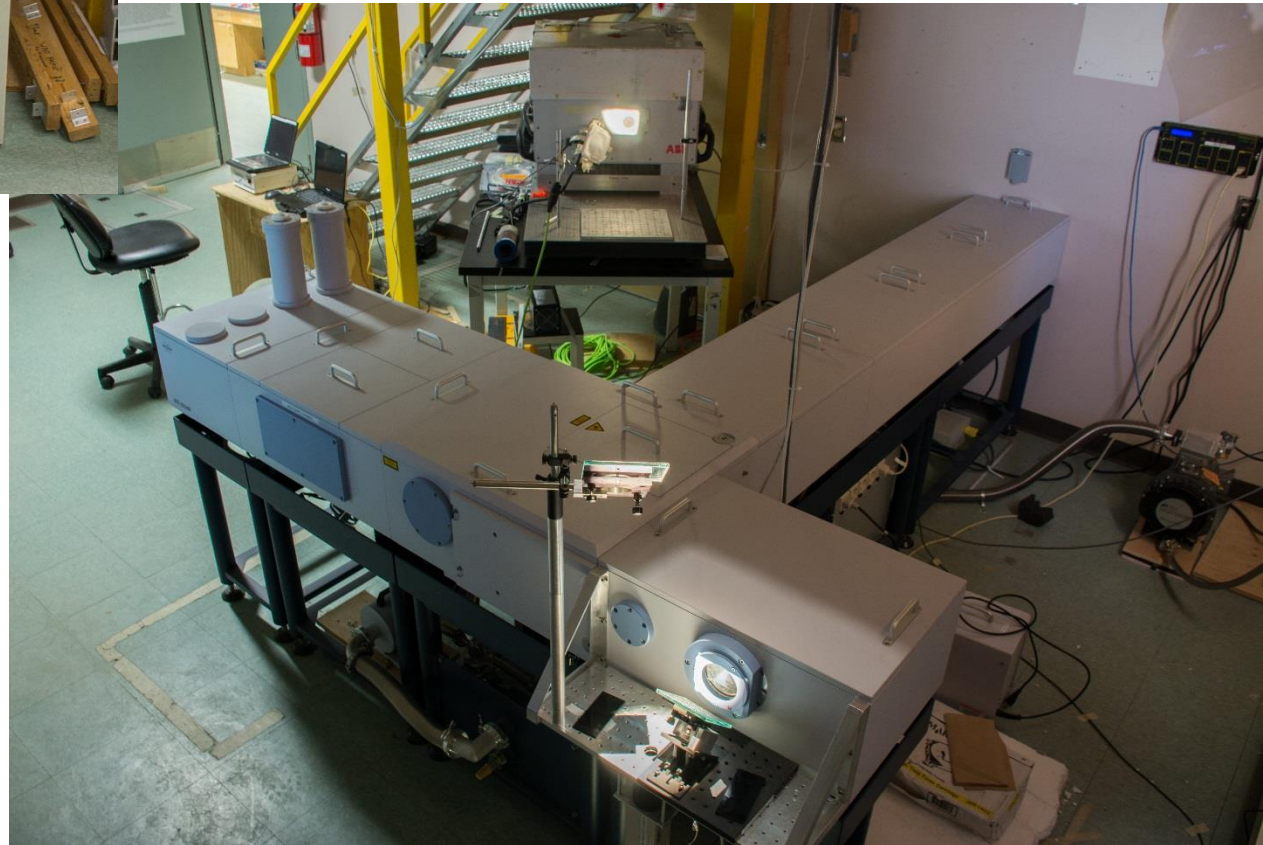
UV-Vis Lab



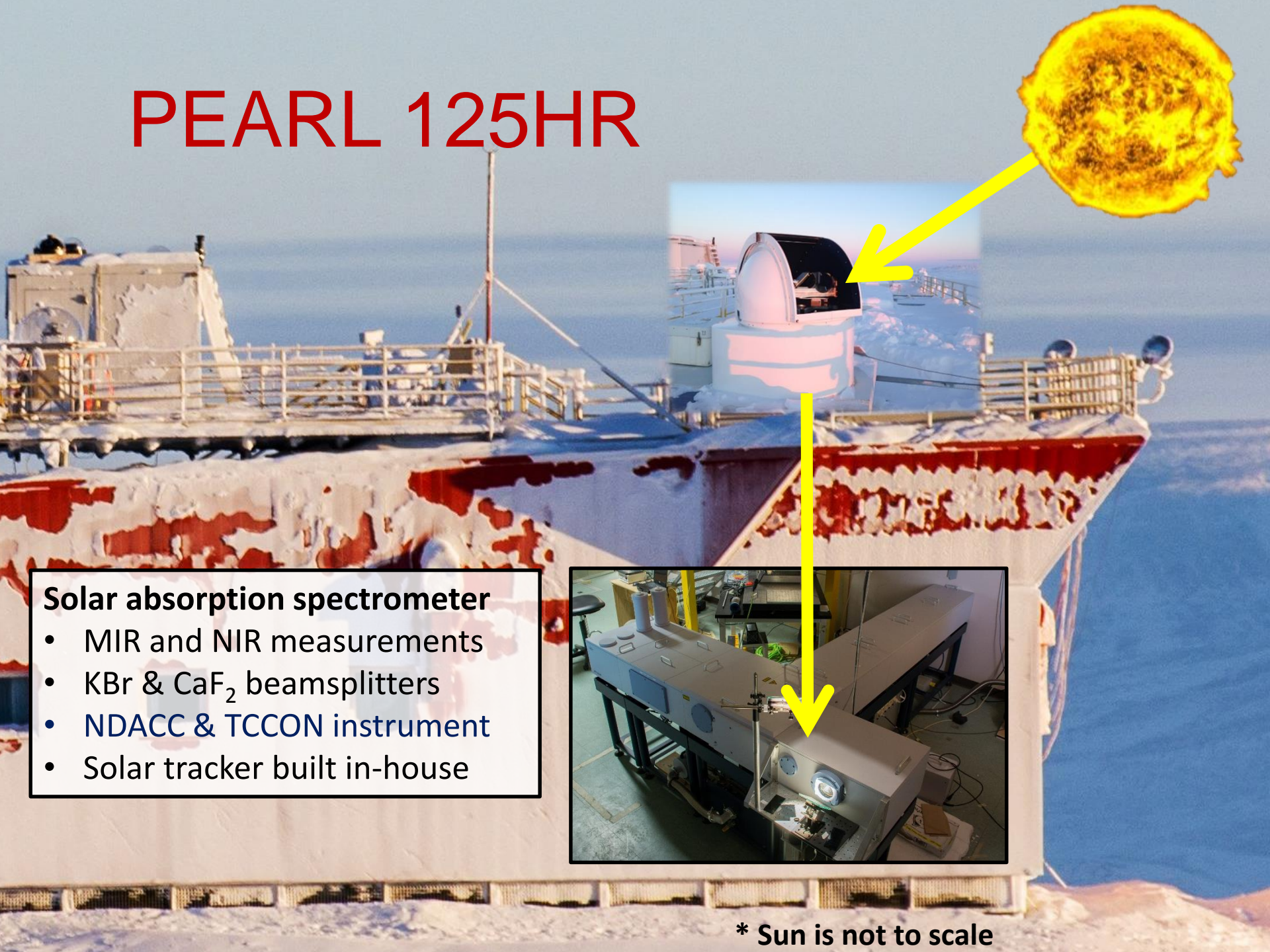
PEARL 125HR



Uses sunlight to
create “maps” of
the atmosphere

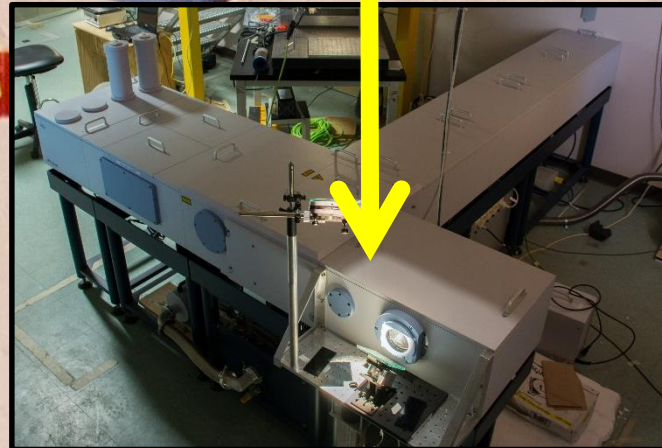


PEARL 125HR



Solar absorption spectrometer

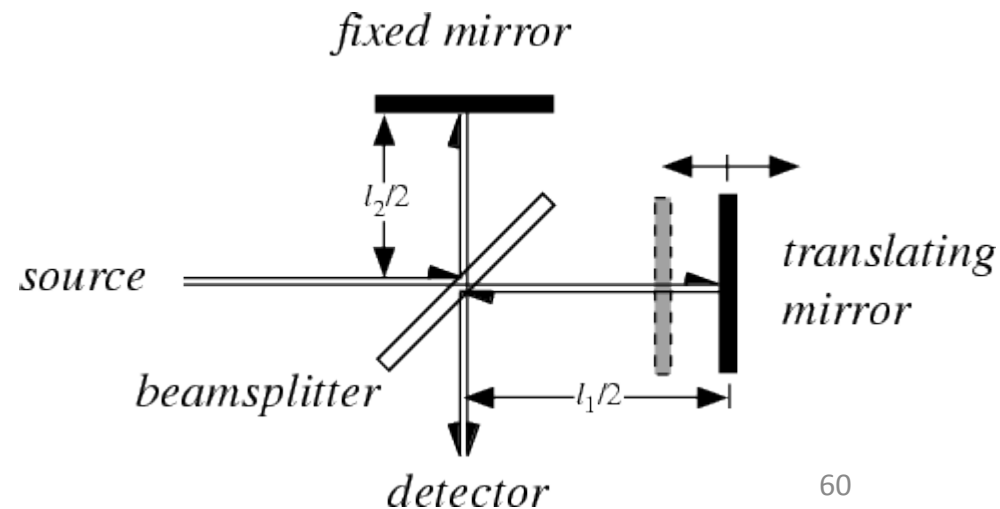
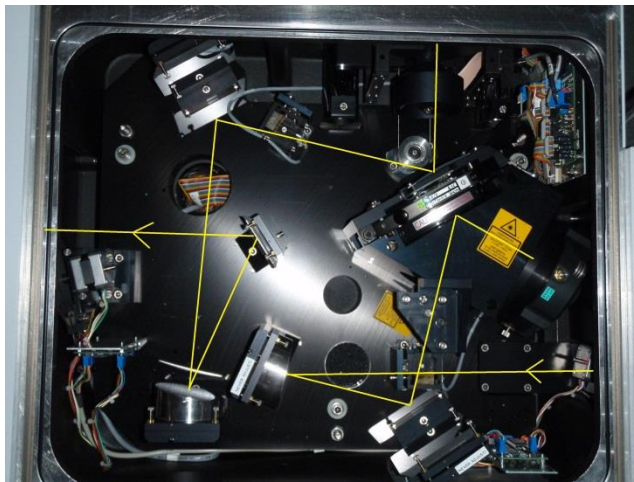
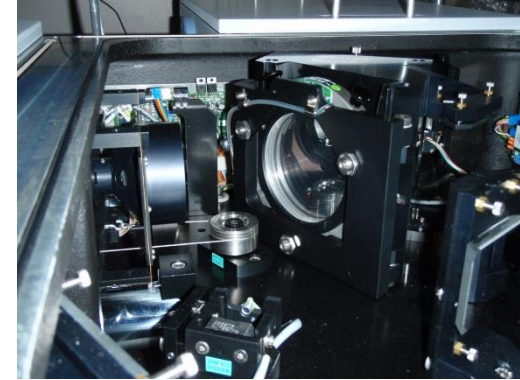
- MIR and NIR measurements
- KBr & CaF₂ beamsplitters
- NDACC & TCCON instrument
- Solar tracker built in-house



* Sun is not to scale

Fourier transform spectroscopy

- Light is passed through a beam splitter
- Two beams of light travel different path lengths
- Detector measures an interference pattern
- As the moving mirror changes position, different frequencies of light peak at different times. The result of a scan is an interferogram.
- The Fourier transform of the interferogram is the observed spectrum, which can be analyzed to retrieve gas abundances and distribution.
- The spectral resolution of an FTS is directly related to the maximum optical path difference

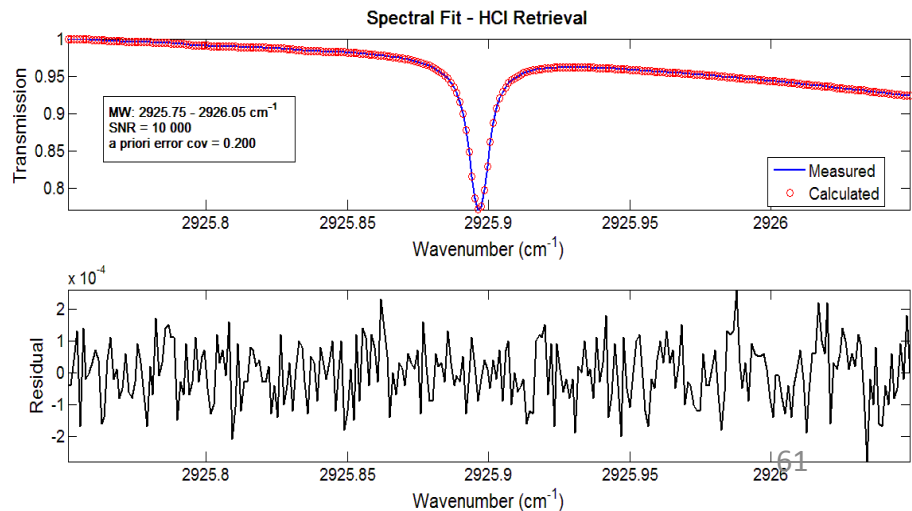


Spectral features

- Extensive lab tests provide a database of spectral features for atmospheric gases.
 - There are 64 000 water vapour lines!
- This knowledge enables us to select appropriate features to examine.
- Inferring abundances and vertical distribution relies on the shape of the observed spectral line.
 - The line strength is directly related to the abundance.
 - The shape is related to the distribution.

$$S = \int_{-\infty}^{\infty} k_{\nu} d\nu$$

where k_{ν} is the absorption coefficient of the molecule being studied



Bruker IFS 125 HR

Detectors

Sample

Interferometer

Moving Mirror
Compartment

Moving Mirror

Beamsplitter

CaF₂
1000 – 25 000 cm⁻¹
KBr
200 – 13 000 cm⁻¹

Fixed Mirror

Glowbar
500 – 4000 cm⁻¹

Tungsten
Bulb
2000 – 25 000 cm⁻¹

Source

InSb
1800 – 11 000 cm⁻¹

MCT
650 – 11 000 cm⁻¹

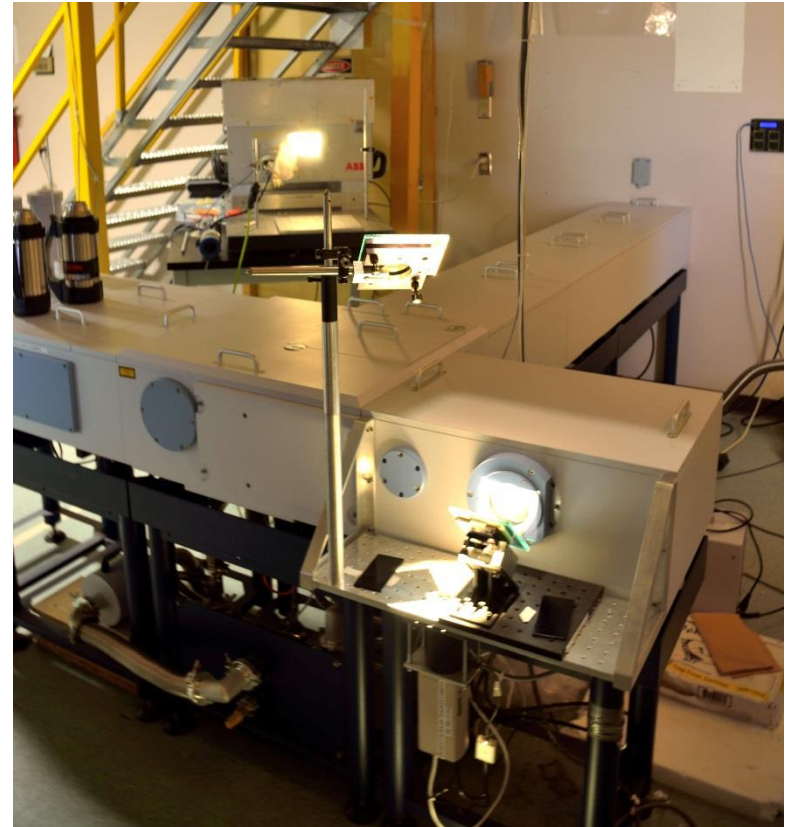
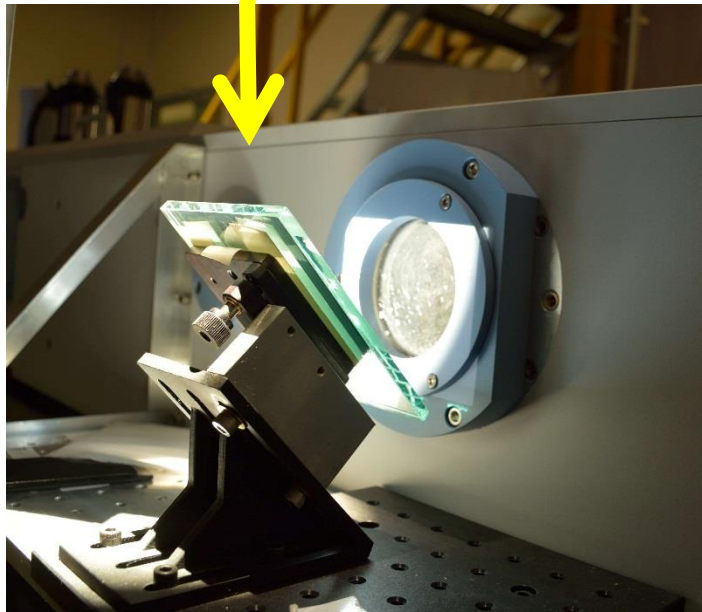
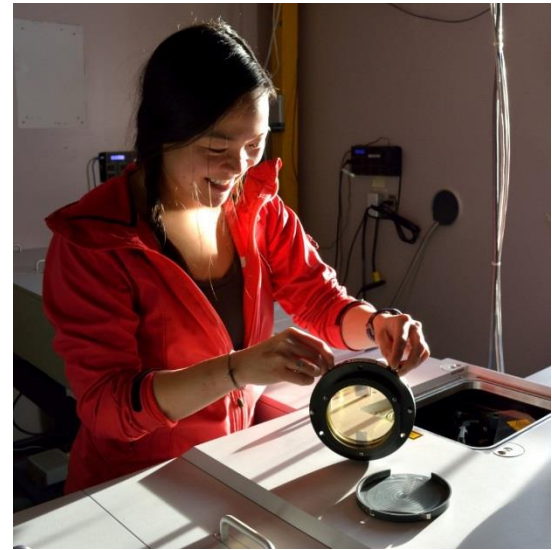
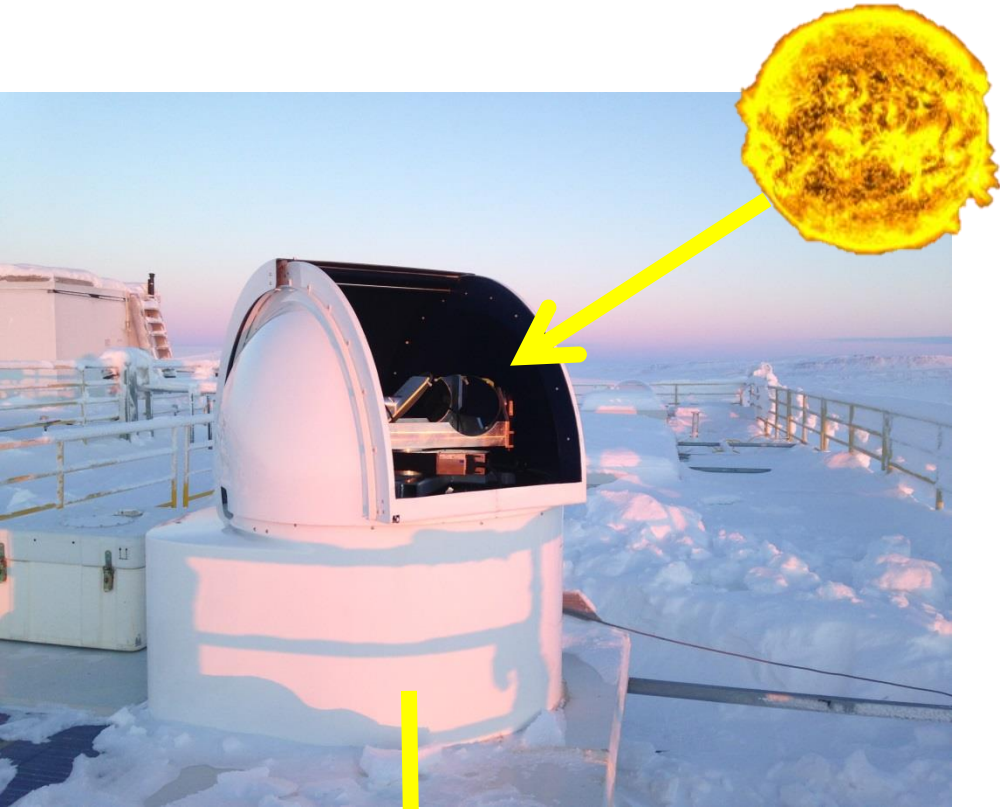


NDACC narrow bandpass filters used at Eureka

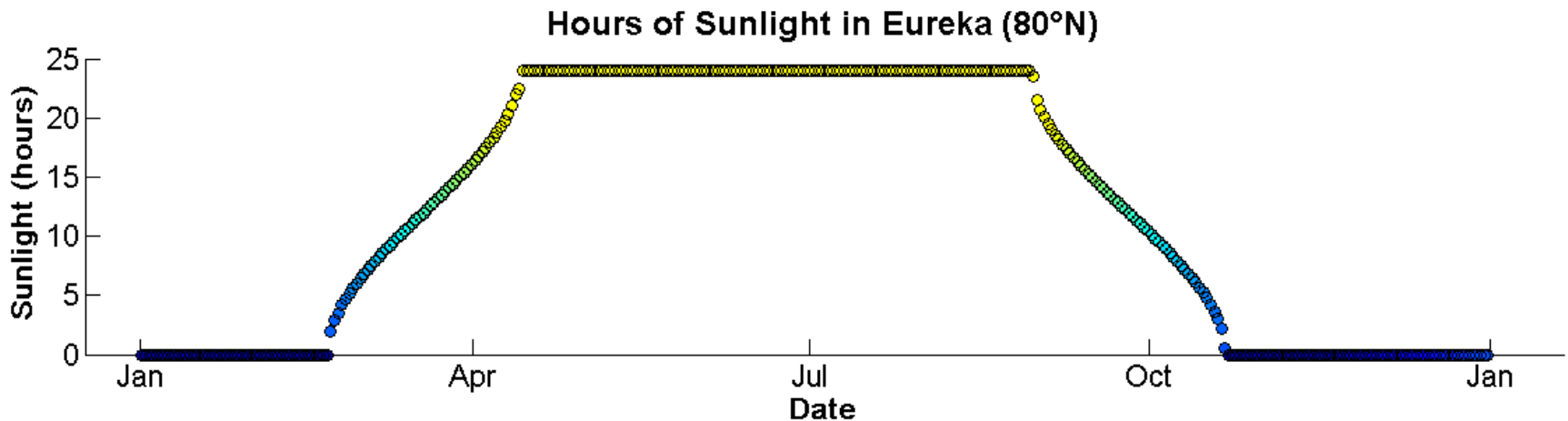
Filter number	Wavenumber range (cm ⁻¹)	Wavelength range (μm)
1	3950 – 4300	2.32 – 2.53
2	2700 – 3500	2.85 – 3.70
3	2420 – 3080	3.25 – 4.13
4	1950 – 2700	3.70 – 5.13
5	1800 – 2200	4.55 – 5.55
6	700 – 1350	7.40 – 14.28
7*	700 – 1000	10.00 – 14.28
8*	1000 – 1400	7.14 – 10.00

* these filters were installed in August 2007.

Table from Lindenmaier (2012)



While a powerful technique, FTIRs are limited to clear weather, sunny conditions.

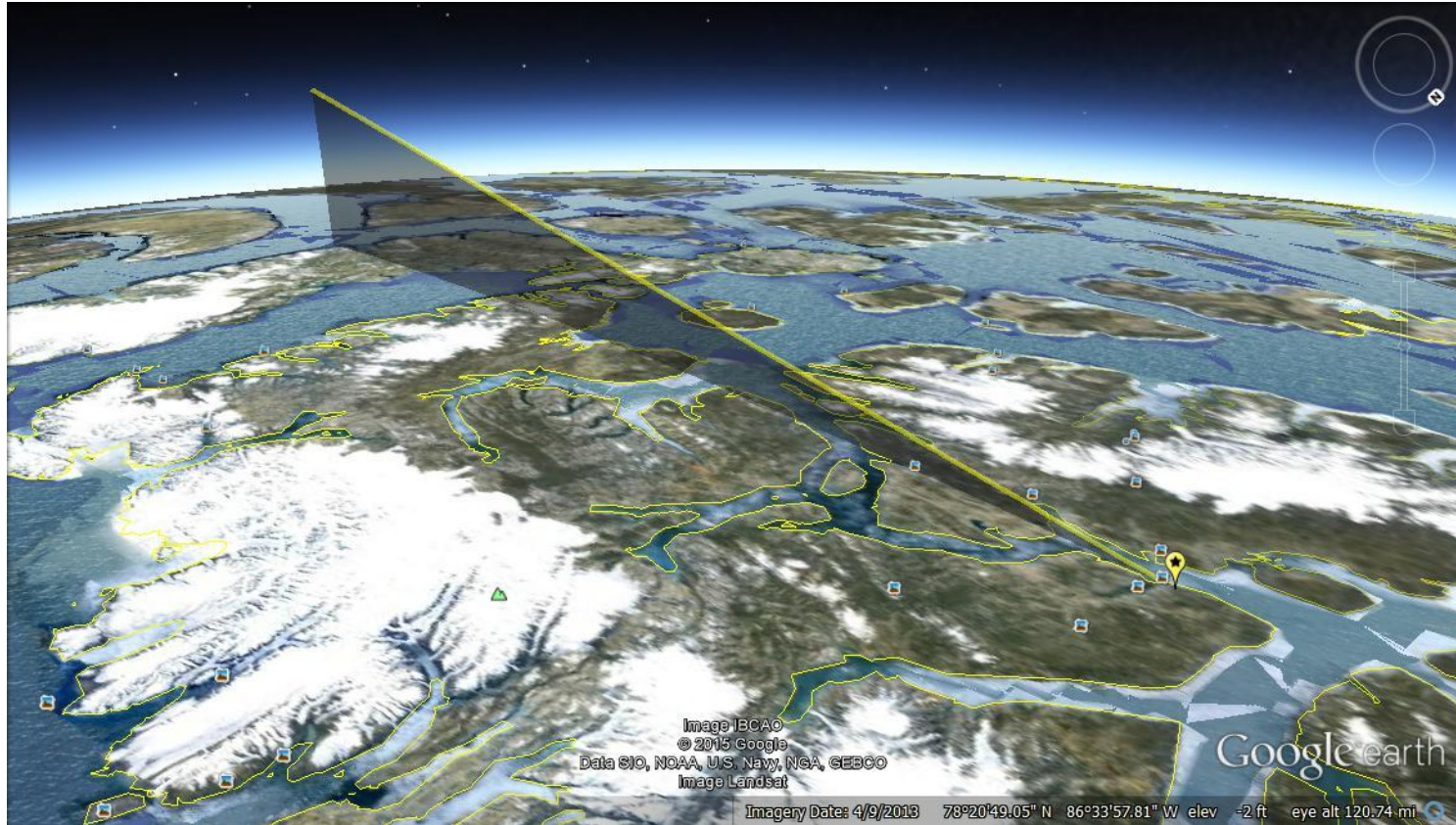


We are currently
experimenting
with using the
moon as a light
source...

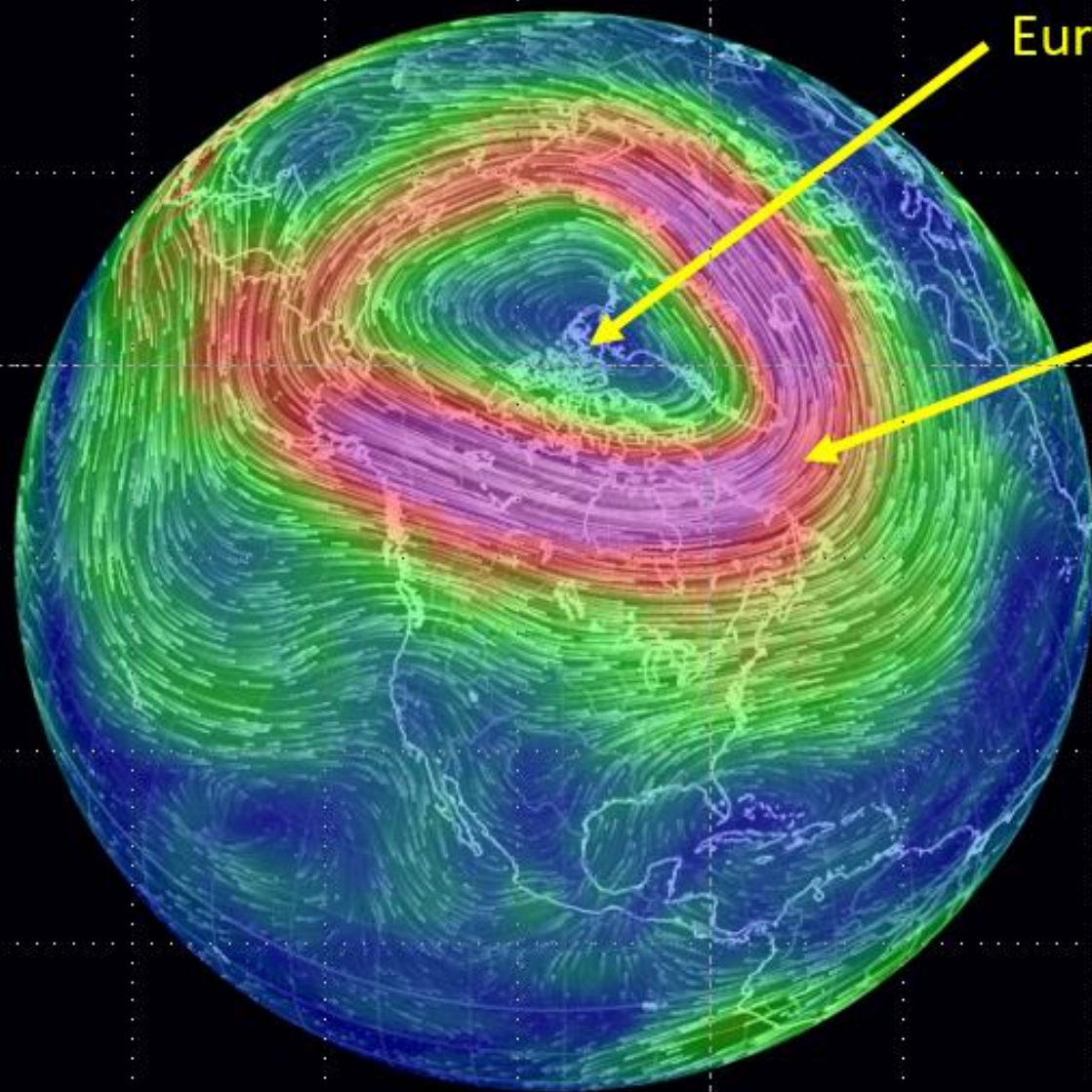


Because the FTIR is looking at the Sun from a high latitude, we're actually looking south

Measurement at 2 pm, March 10, 2014



At 8 km altitude, we're measuring ~ 70km away
At 20 km altitude, we're measuring ~150 km away
At 35 km altitude, we're measuring ~250 km away
At 50 km altitude, we're measuring ~350 km away
At 100 km altitude, we're measuring ~600 km away

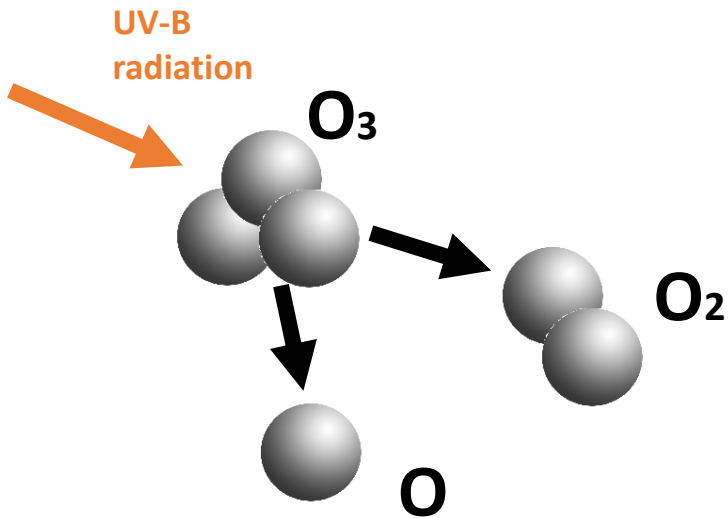


Eureka/PEARL

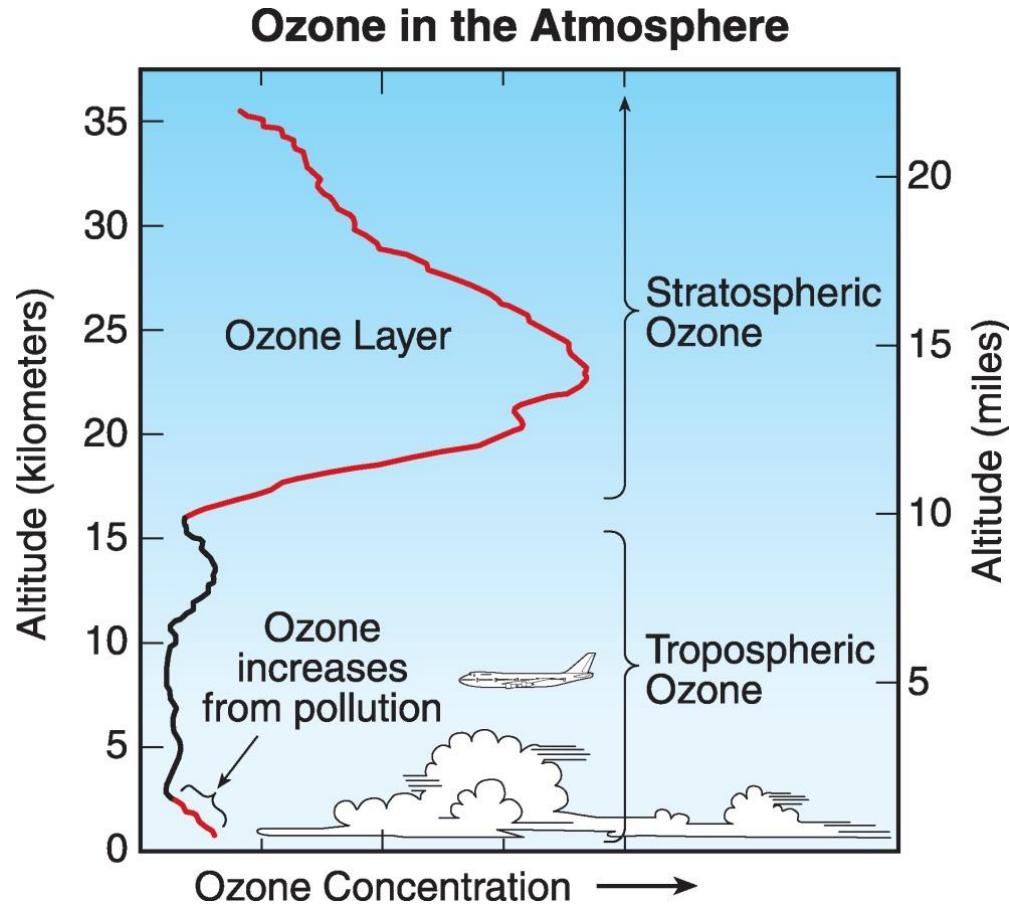
Polar Vortex

What are we using
these measurements
to study?

Ozone



- Ozone in the stratosphere protects us from harmful UV radiation (good!)
- Ozone in the troposphere is a pollutant and harmful to breath! (bad!)

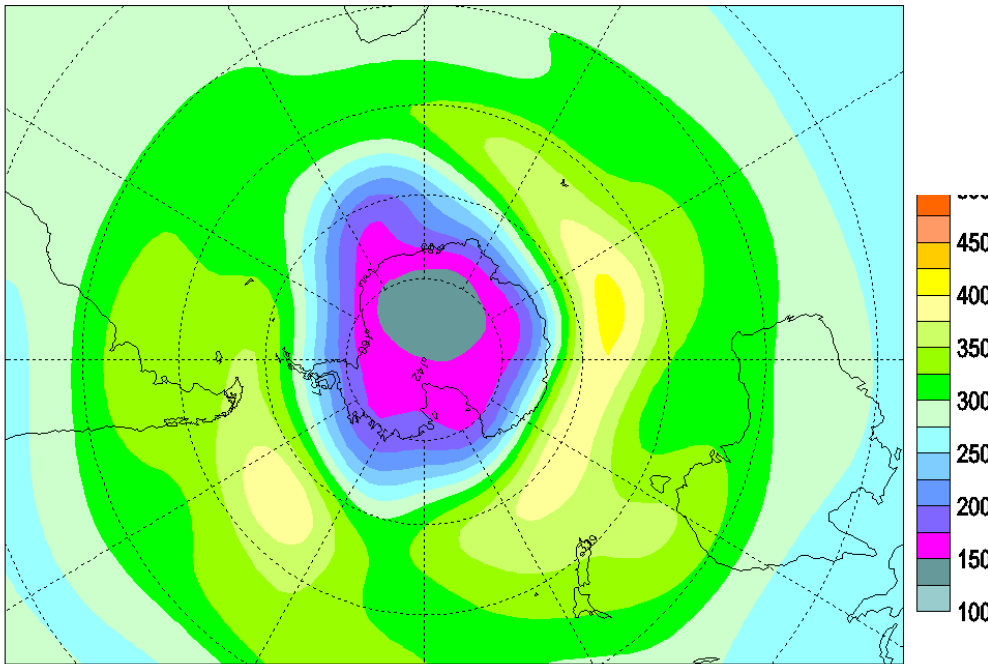


UNEP/WMO, 2006: Fig. Q1-2

“Space isn't remote at all. It's only an hour's drive away if your car could go straight upwards.” – Sir Frederick Hoyle

Ozone 'Hole' in Antarctic: First Discovered in 1985

Mean total ozone (DU), 2010/10/01-2010/10/31

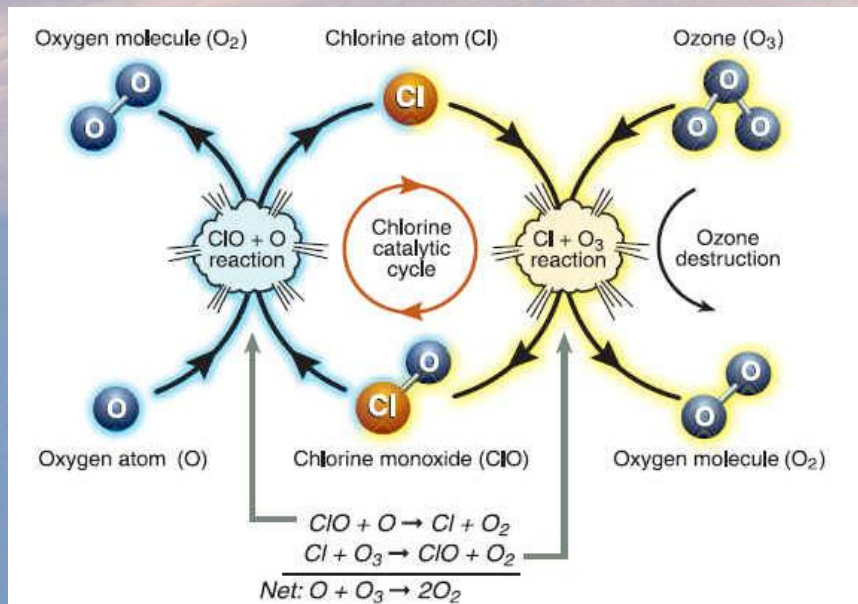


What is an Ozone 'Hole'?

- When there is very *little* ozone
- ~ 1/3 (or less) than usual
- Puts plants and animals at greater risk of UV radiation exposure.
- Too much UV radiation can damage skin cells, which can lead to skin cancer and strong UV radiation can damage the cell structure of plant material.

How CFCs destroy ozone

- When CFCs reach the ozone layer, ultraviolet sunlight breaks them apart to form atomic chlorine.

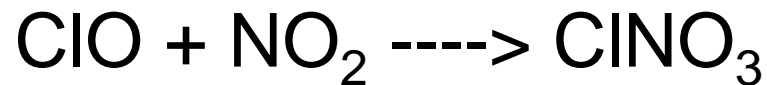
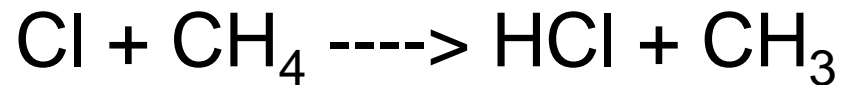


One chlorine atom can destroy thousands of ozone molecules, because it is “recycled” through catalytic cycles.

Ending Ozone Destruction

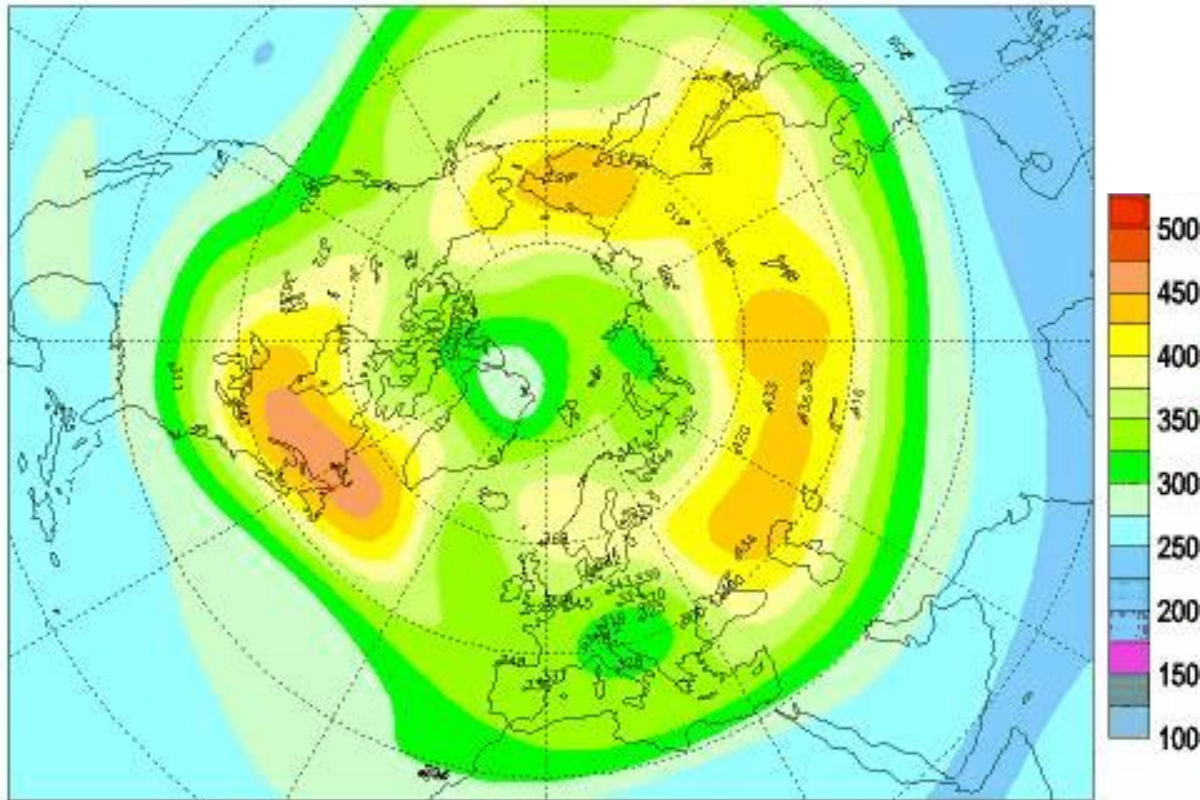
How does this **stop**?

- When Cl encounters methane, CH₄ or when ClO encounters NO₂, the cycle terminates



A Big Surprise: 2011 Arctic Ozone Hole

Total ozone (DU) / Ozone total (UD), 2011/02/10



This isn't good...

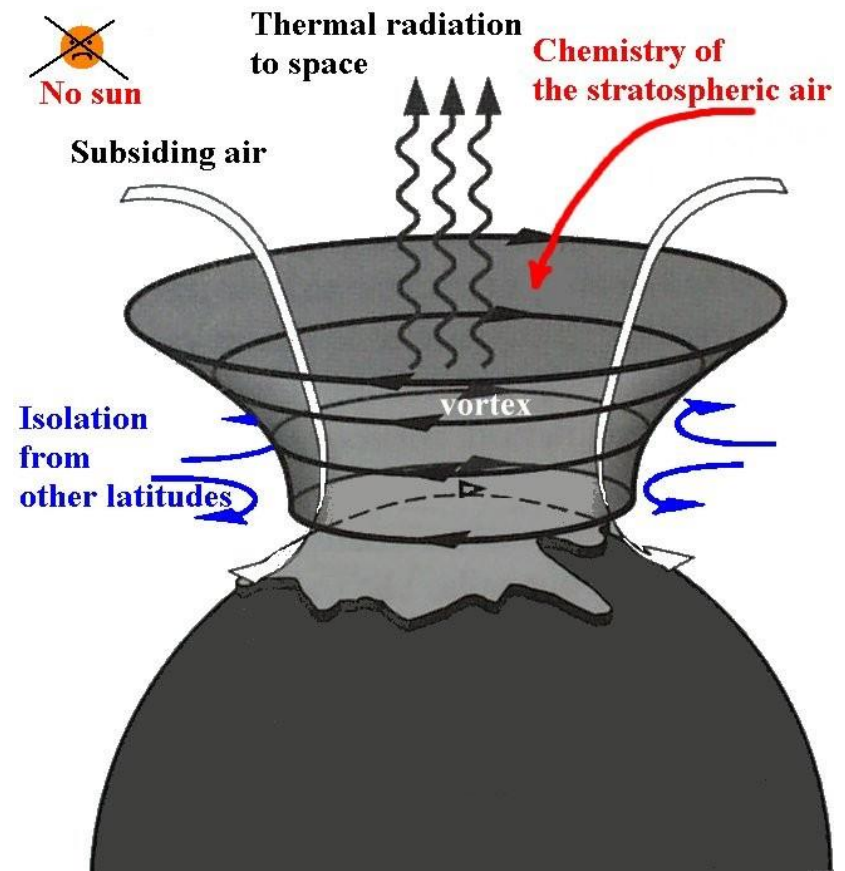


Colder than usual temperatures in the Arctic stratosphere caused the largest ozone depletion event yet observed. Debate surrounds use of the term 'hole'.

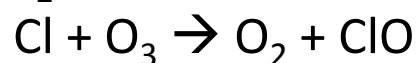
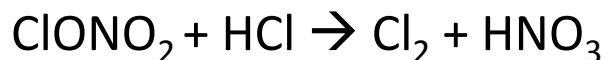
Why Ozone Holes in the polar regions?

The Polar Vortex enables the stratosphere to reach very low temperatures ($< 78^{\circ}\text{C}$), allowing Polar Stratospheric Clouds to form.

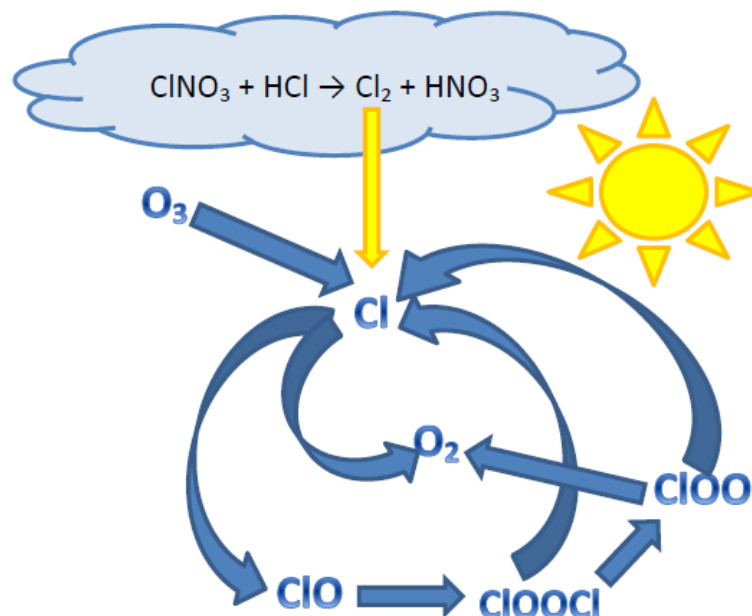
PSCs provide a surface for ozone-destroying chemical reactions to happen.



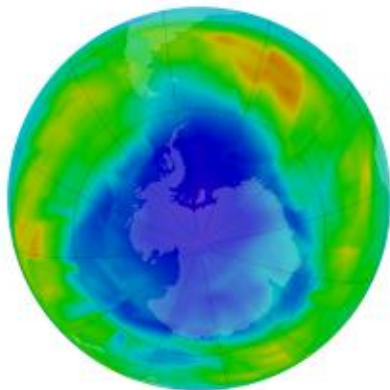
PSC chemistry



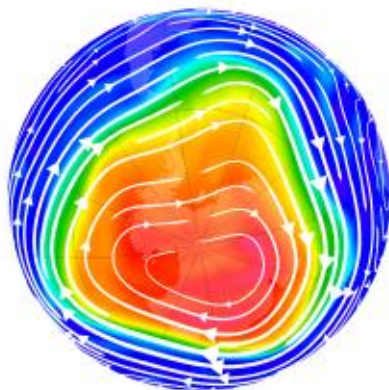
Without PSCs, ClONO_2 and HCl would not generate active chlorine, but instead would remain in a reservoir until eventually deposited to the ground.



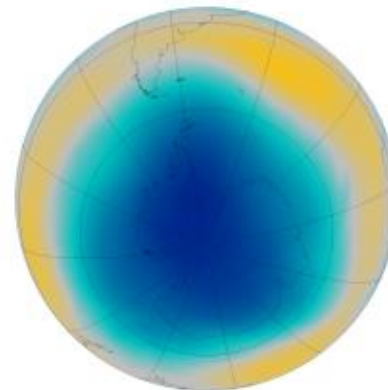
Ozone



Potential Vorticity



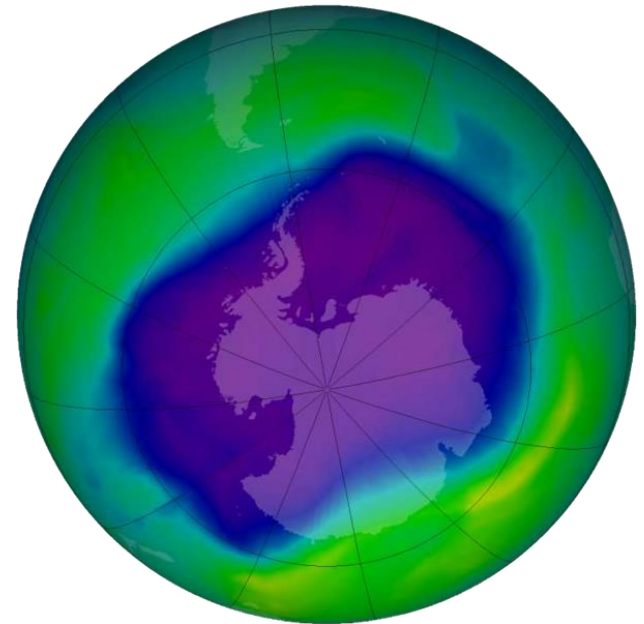
Temperature



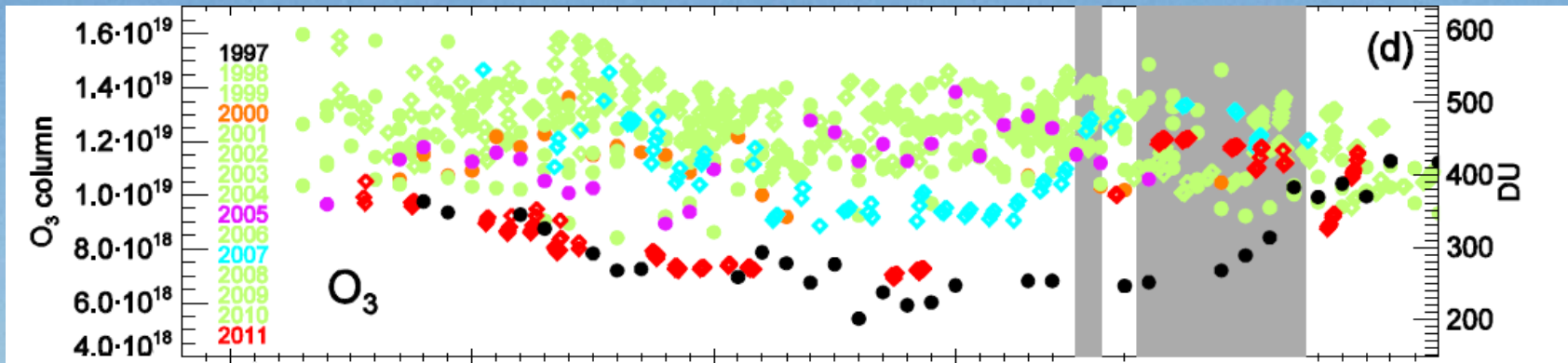
Why is ozone depletion *still* a problem?

- CFCs have a long lifetime.
- It will take at least another 50 years for a full recovery
- This should be a warning for us to be cautious about modifying the Earth's atmosphere...

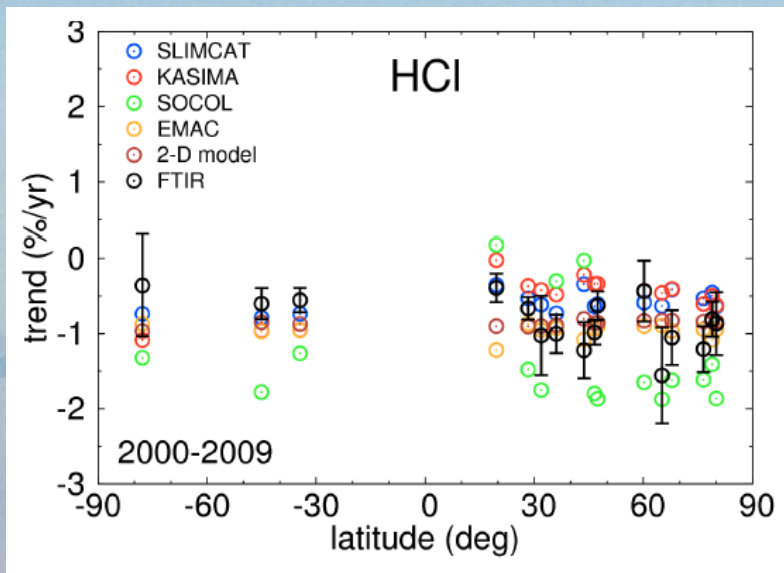
2006's Antarctic ozone hole:



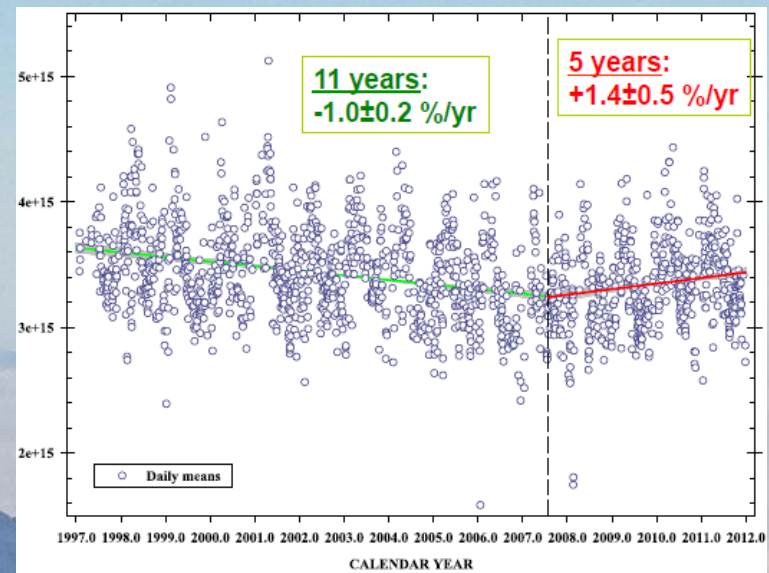
Current research into trends



PEARL O₃ timeseries from Lindenmaier (2012)



Global HCl trends from Kohlhepp (2012)



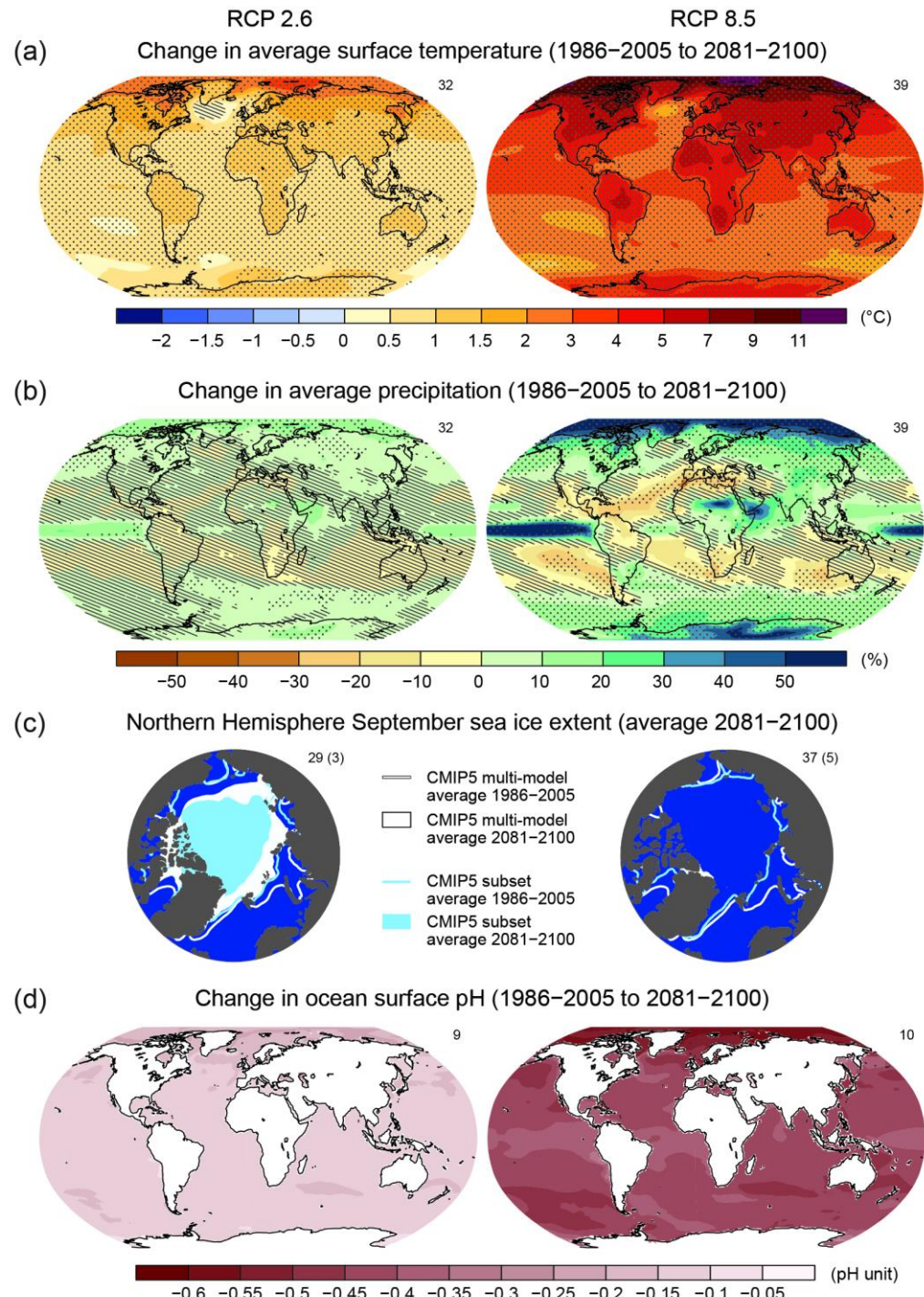
Jungfraujoch HCl from Mahieu (2012)

FTIRs have a useful advantage

- They can take large spectra from which *many* atmospheric gases can be measured simultaneously
- This is a powerful means of observing chemistry.
- PEARL's FTIR measurements are being used to investigate conversions between active and reservoir forms of Chlorine, and the relationships with nitrogen chemistry

Climate change

- Over the last 50 years, the Arctic has warmed by 5°C
(Lesins et al. 2010)
- The Arctic is expected to warm more than & faster than any other area of the world. (IPCC, 2014)

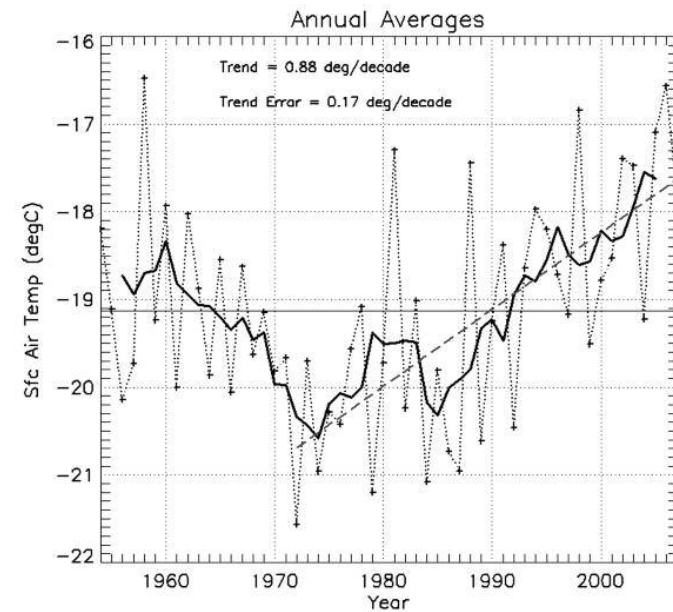
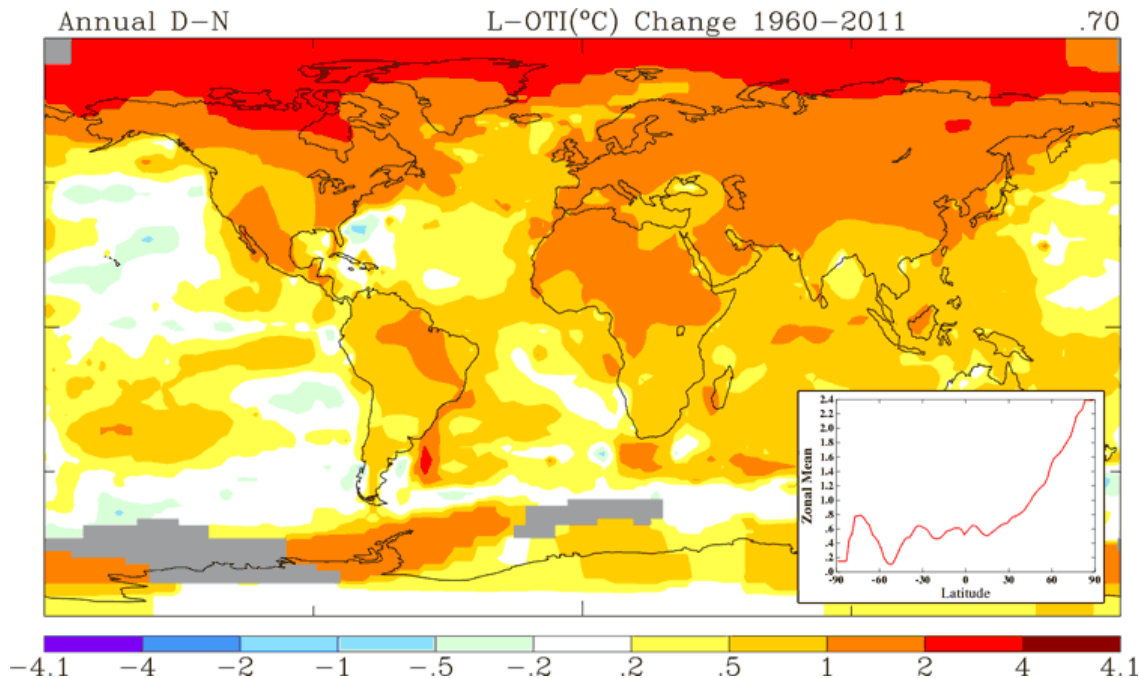


Arctic temperature change

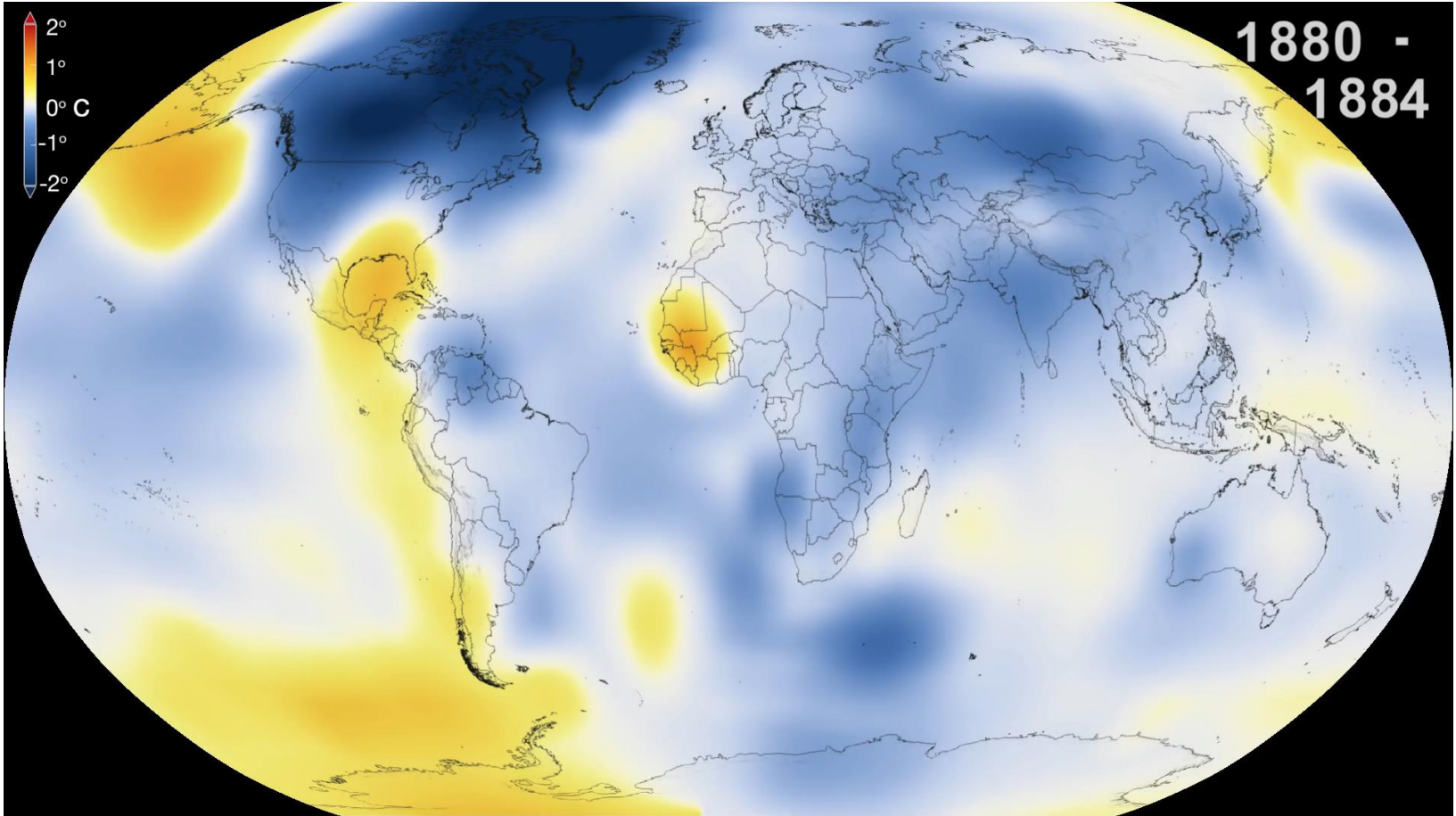
Eureka temperature trend:
0.88°C/decade since 1972
(Lesins et al. 2010)

Total warming:

- 1.5°C summer (least)
- 4.5°C autumn (most)



Global temperature change



Total Carbon Column Observing Network (TCCON)



Arctic climate & water vapour

Mid-latitudes

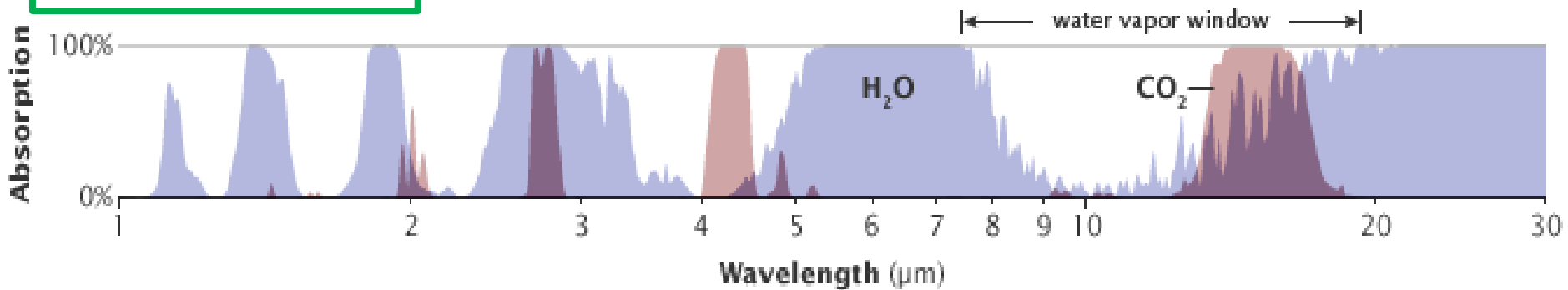


Image credit: <http://earthobservatory.nasa.gov/Features/EnergyBalance/page7.php>

Arctic

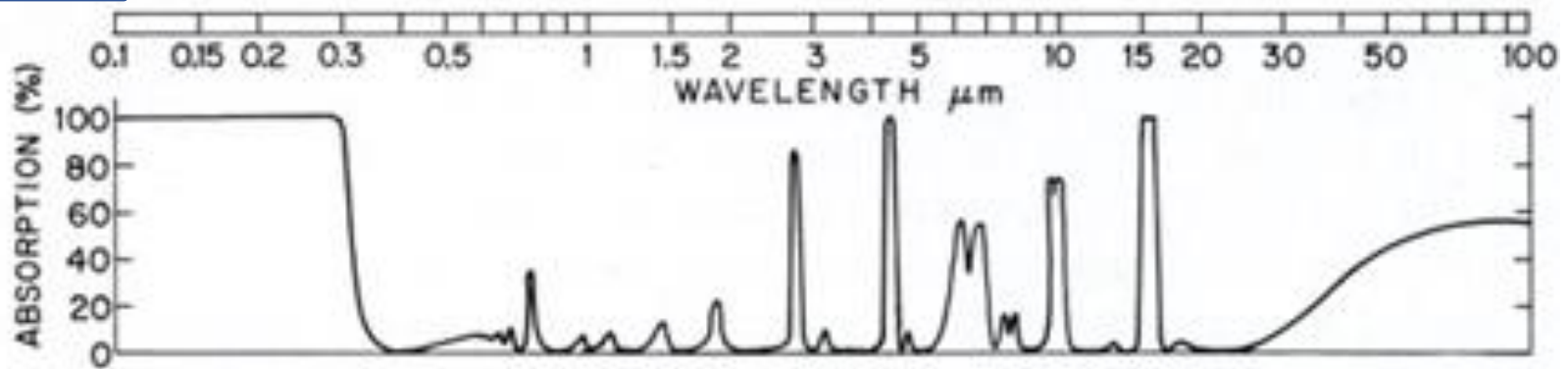
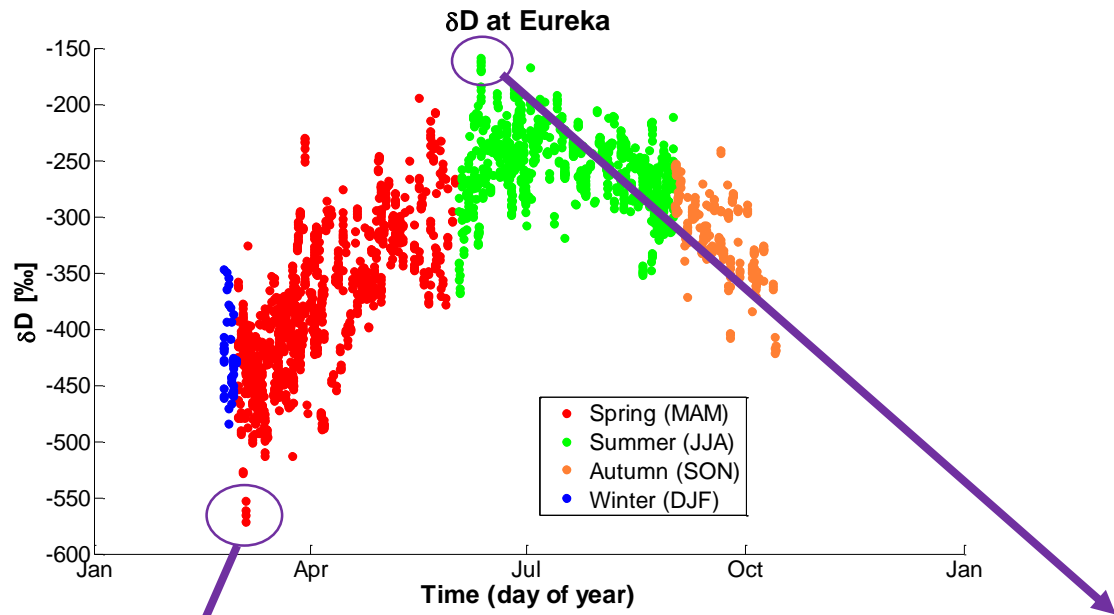
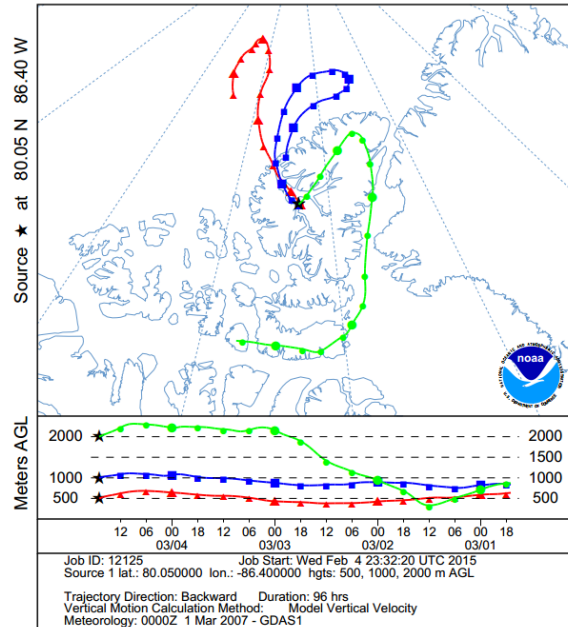


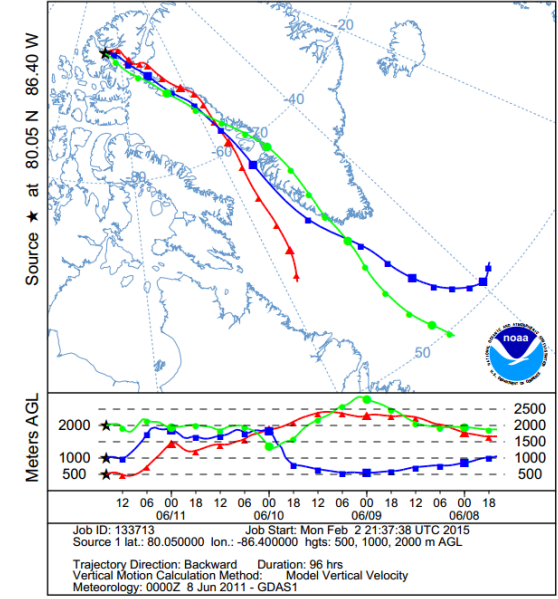
Image credit: Zak, B. and K. Stamnes (1997): ARM Science Meeting – Session Papers. 19-25



NOAA HYSPLIT MODEL
Backward trajectories ending at 1700 UTC 04 Mar 07
GDAS Meteorological Data



NOAA HYSPLIT MODEL
Backward trajectories ending at 1600 UTC 11 Jun 11
GDAS Meteorological Data

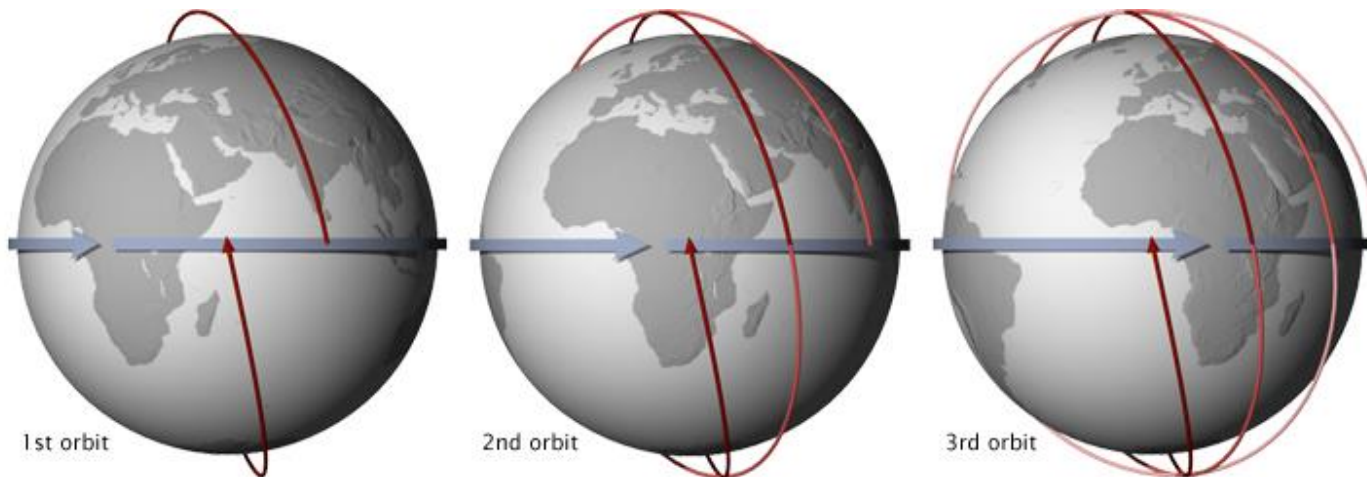


One more science objective...

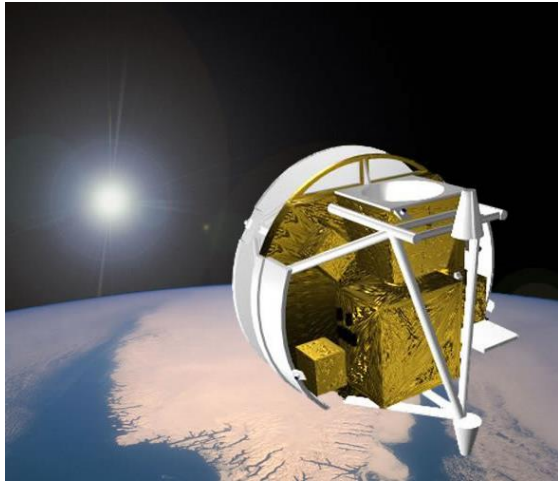
Satellite validation

Eureka is easy to get to ...for a satellite!

Sun-synchronous orbits provide global coverage,
and frequently pass over the poles



ACE (Canada)



OCO-2 (U.S. NASA)



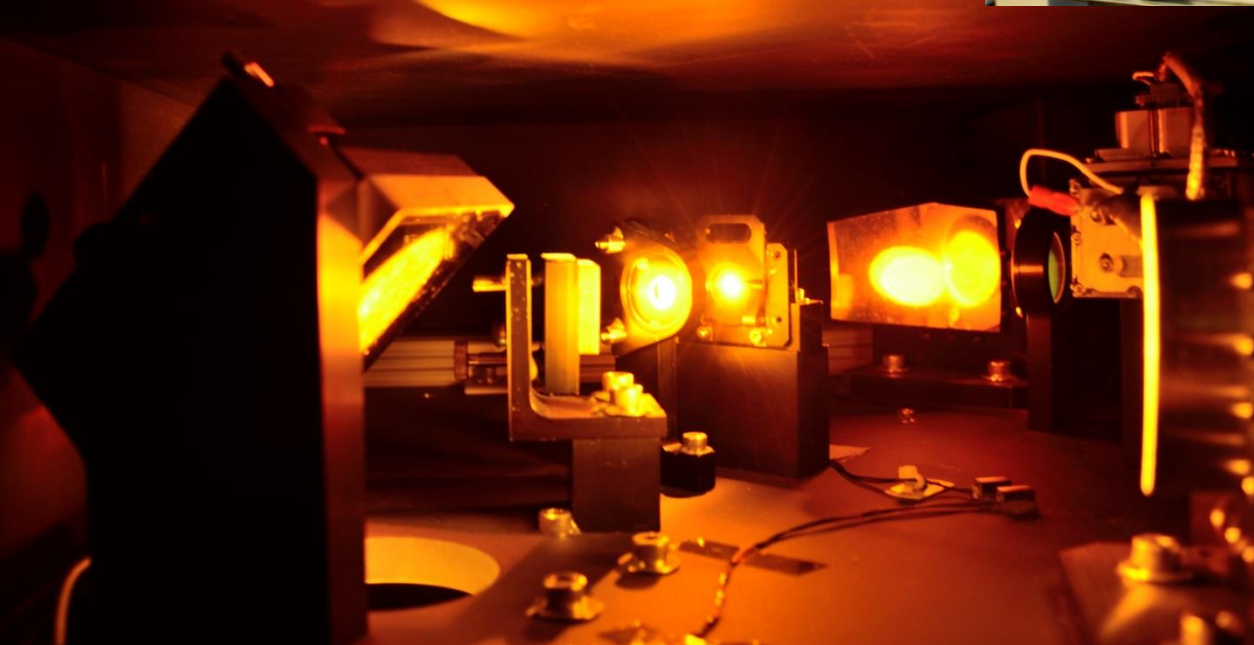
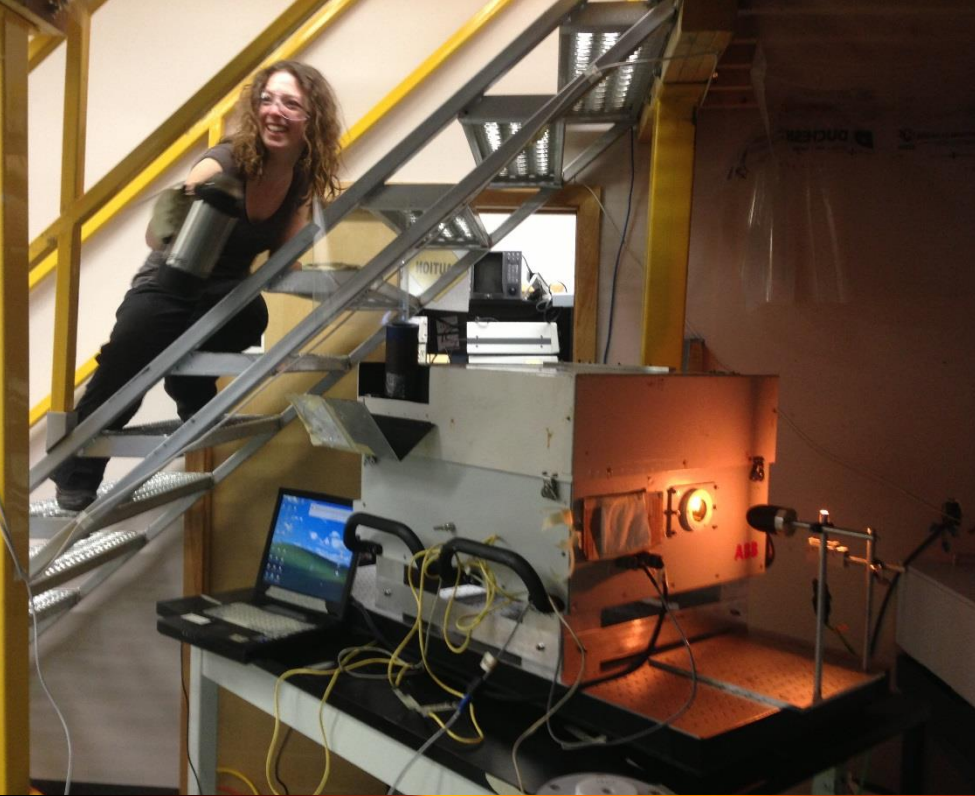
Odin (Sweden & Canada)



GOSAT (Japan)



PARIS



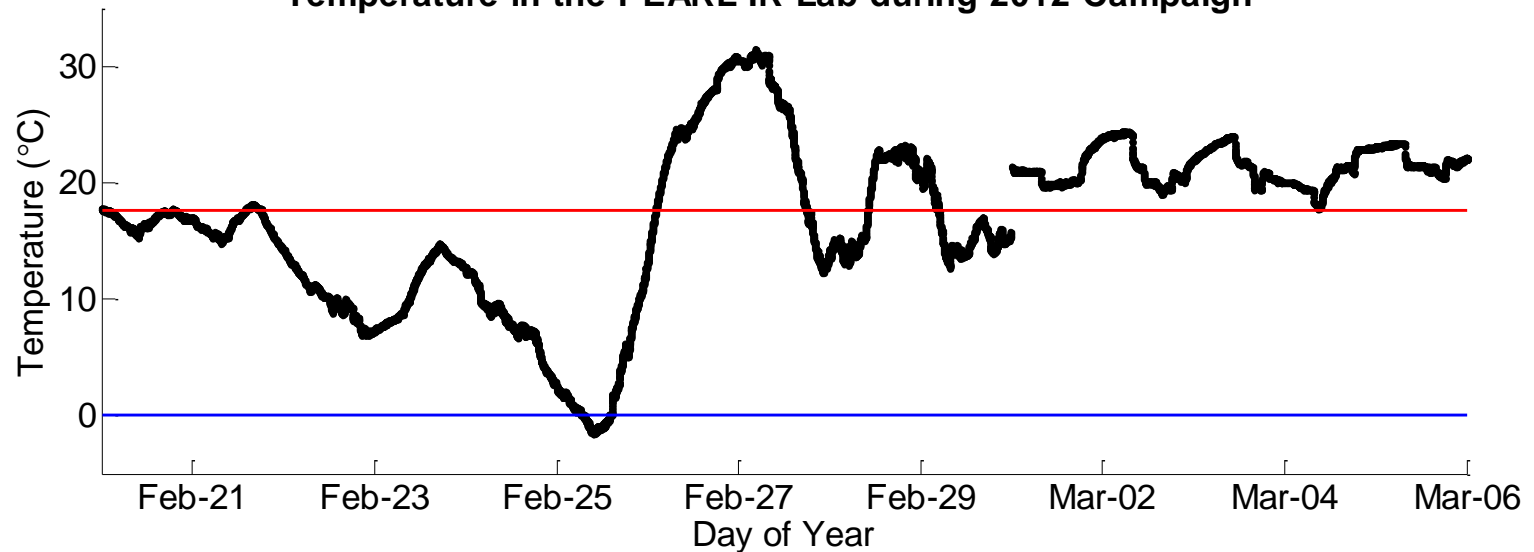
Working in the extreme cold is challenging, for people & equipment



Operational challenges abound



Temperature in the PEARL IR Lab during 2012 Campaign





© Dan Weaver

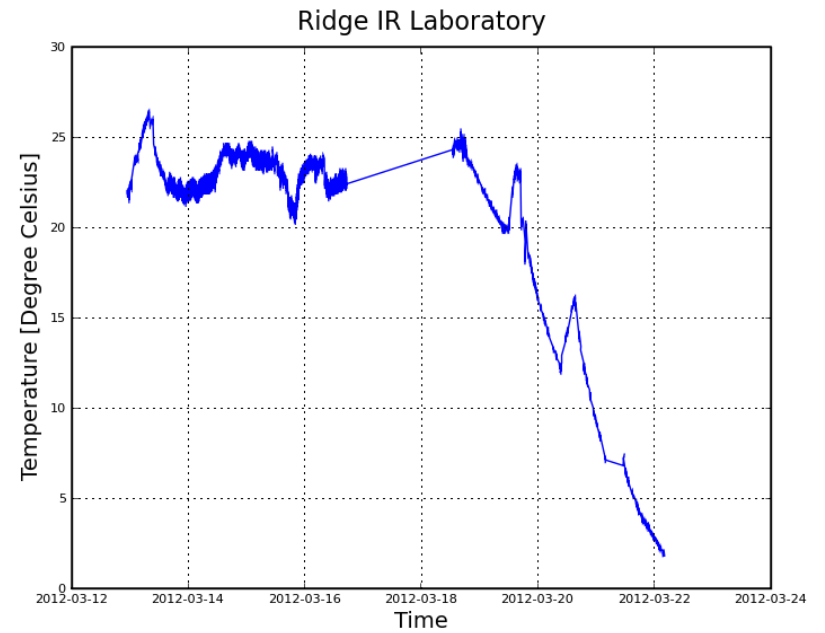
Upgrades



Heat & Cooling



Arctic challenges



Internet

- Website for CANDAC, which runs PEARL
- Campaign website
- Social media
 - Twitter
 - Instagram
 - blog



Instagram profile for **createarcticsci**. The profile picture shows a person in a snowy landscape. The bio reads: "Run by @Dan rhythm for CREATE program in Arctic atmospheric science. We include researchers at many Canadian universities. Often working at PEARL. <http://createarcticscience.wordpress.com>". The statistics show 36 posts, 6 followers, and 4 following. The background features a grid of six images: a ship on ice, a person in a snowfield, a person in a lab, a person in a snowfield, a person in a lab, and a person in a snowfield.



Twitter profile for **CREATE ArcticScience** (@CREATEArcticSci). The profile picture shows a person in a snowy landscape. The bio reads: "Run by the Trainee Advisory Committee of the NSERC CREATE Training Program in Arctic Atmospheric Science. We include researchers at many Canadian universities. Canada · createarcticscience.wordpress.com". The statistics show 307 tweets, 149 following, and 189 followers. A tweet from **U of T Magazine** (@uoftmagazine) is shown, dated Mar 20, with the text: "Why a freezing fogbank was good news for Canada's farthest-North scientists @CREATEArcticSci bit.ly/1gKzhAq pic.twitter.com/14v3YZ5Qxk". The background features a snowy landscape with a sunset.

Twitter: jokes with NASA engineer & former internet meme


 **Bobak Ferdowsi** @tweetsoutloud · Feb 27
Is this the rebel base on Hoth? MT"@csa_asc: @createarcticsci: 1st day at PEARL pic.twitter.com/NSP1P0iyde"







 Hide photo  Reply  Retweeted  Favorited  More  HootSuite




RETWEETS: 66 FAVORITES: 122

6:46 PM - 27 Feb 2014 · Details Flag media

 **Bobak Ferdowsi** @tweetsoutloud · Mar 1
Follow up: ion cannon at Hoth. MT"@CREATEArcticSci: we have a powerful laser at PEARL! pic.twitter.com/AgWOrzowE4"

 Hide photo  Reply  Retweeted  Favorited  More  HootSuite



RETWEETS: 30 FAVORITES: 52

1:30 PM - 1 Mar 2014 · Details Flag media

& more fun



Thank you.





More information

www.twitter.com/CREATEarcticsci

www.instagram.com/CREATEarcticsci

<http://createarcticscience.wordpress.com/>

<http://acebox.uwaterloo.ca/eureka/>

CANDAC.ca

Contact me

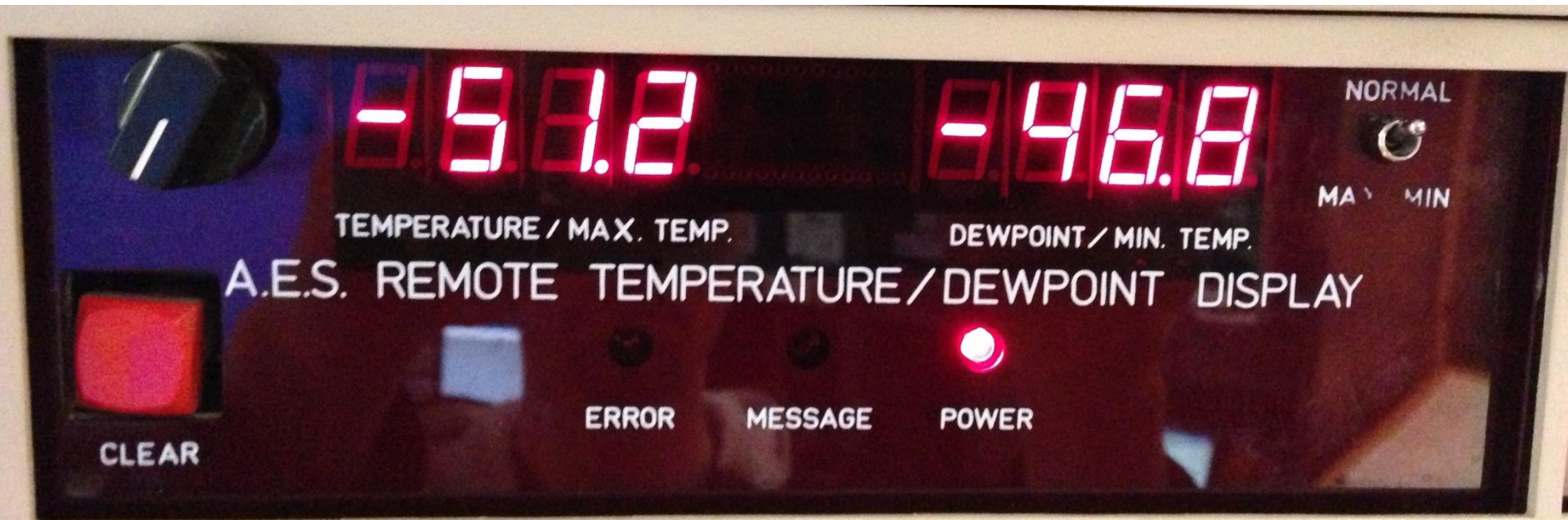
dweaver@atmosp.physics.utoronto.ca

A person is shown in silhouette, standing in a dark room and looking out a window. The person's hands are raised to their forehead, suggesting a moment of discovery or contemplation. The window is partially covered by light-colored curtains, and bright light streams in from outside. A text box with a blue border is overlaid on the right side of the image, containing the text "Exploring Eureka" in red. In the top left corner, there is a small logo with the text "2006" and "V".


**Exploring
Eureka**



Temperature

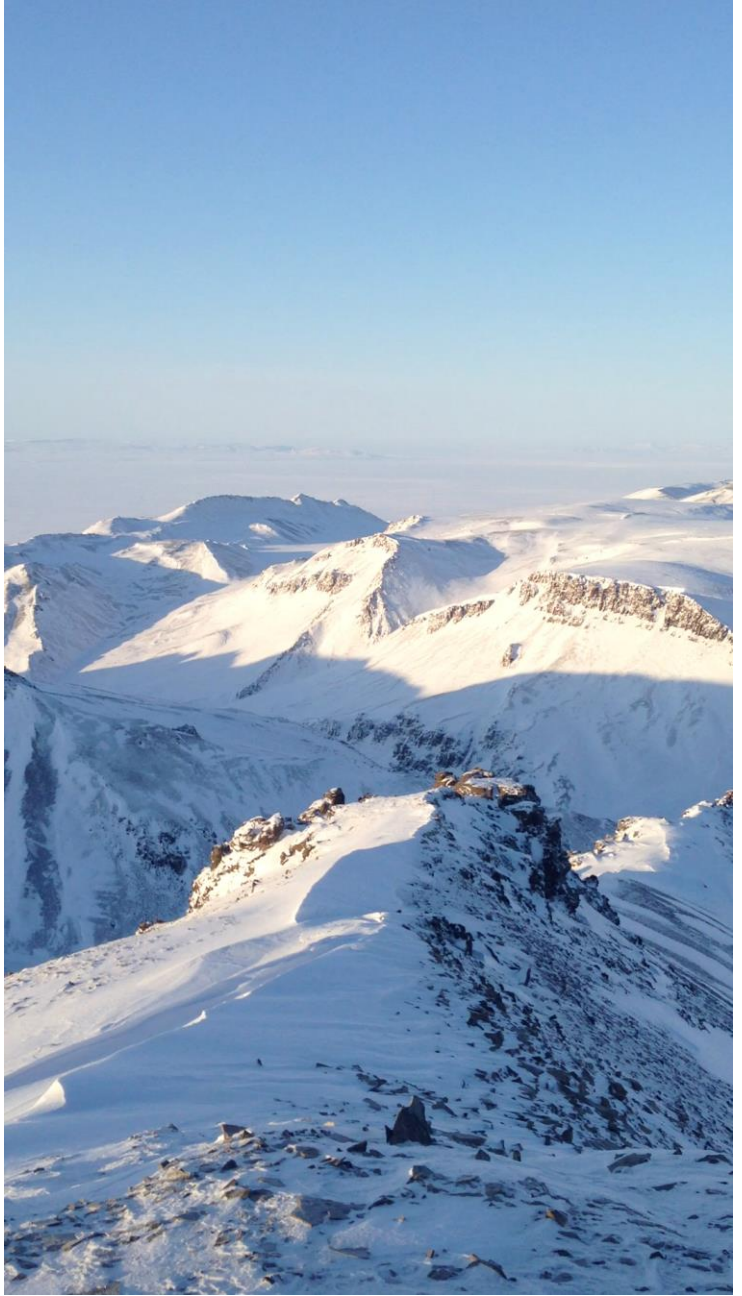


Dress warmly!



**Bring a sense of
adventure &
curiosity**









Night hikes





Arctic Wildlife







© Dan Weaver







© Dan Weaver







© Dan Weaver



© Dan Weaver



© Dan Weaver











Scenery





© Dan Weaver





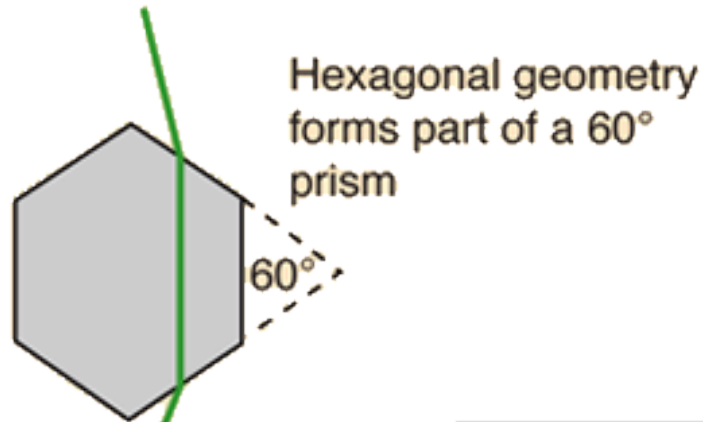


Parhelia, or “sun dogs” (parhelion is singular)



Halo & sundog explanation

Hexagonal ice crystals refract the sun's rays to give the familiar 22° halo.



The order of colors is reversed from that of diffraction, and more closely spaced.

660 nm	550 nm	410 nm
$n=1.306$	$n=1.311$	$n=1.317$
21.54°	21.92°	22.37°
red	green	blue









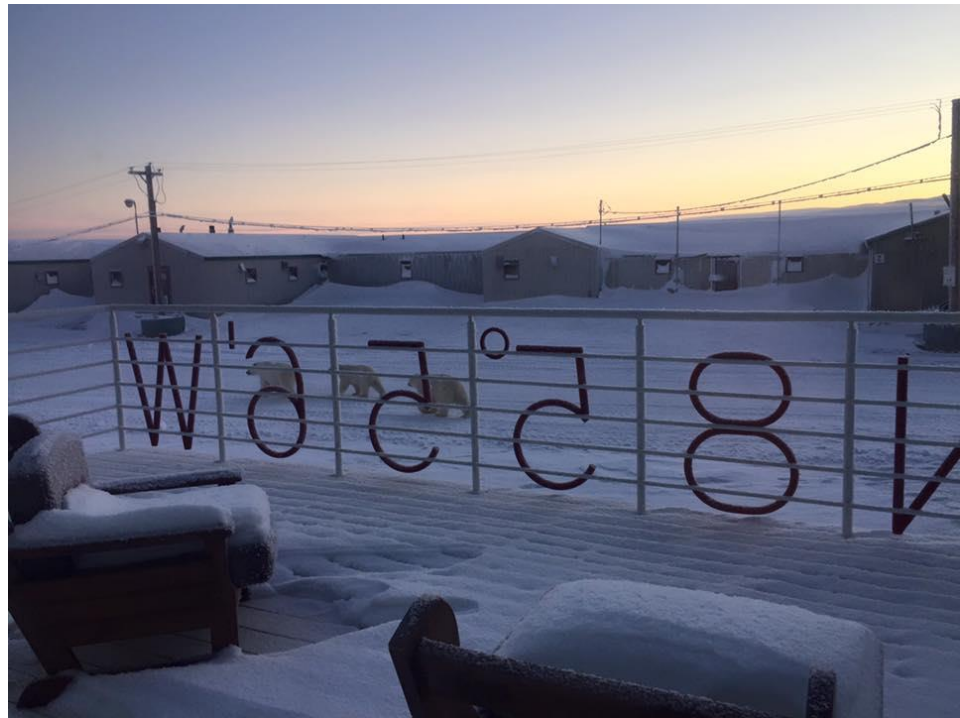


© Dan Weaver

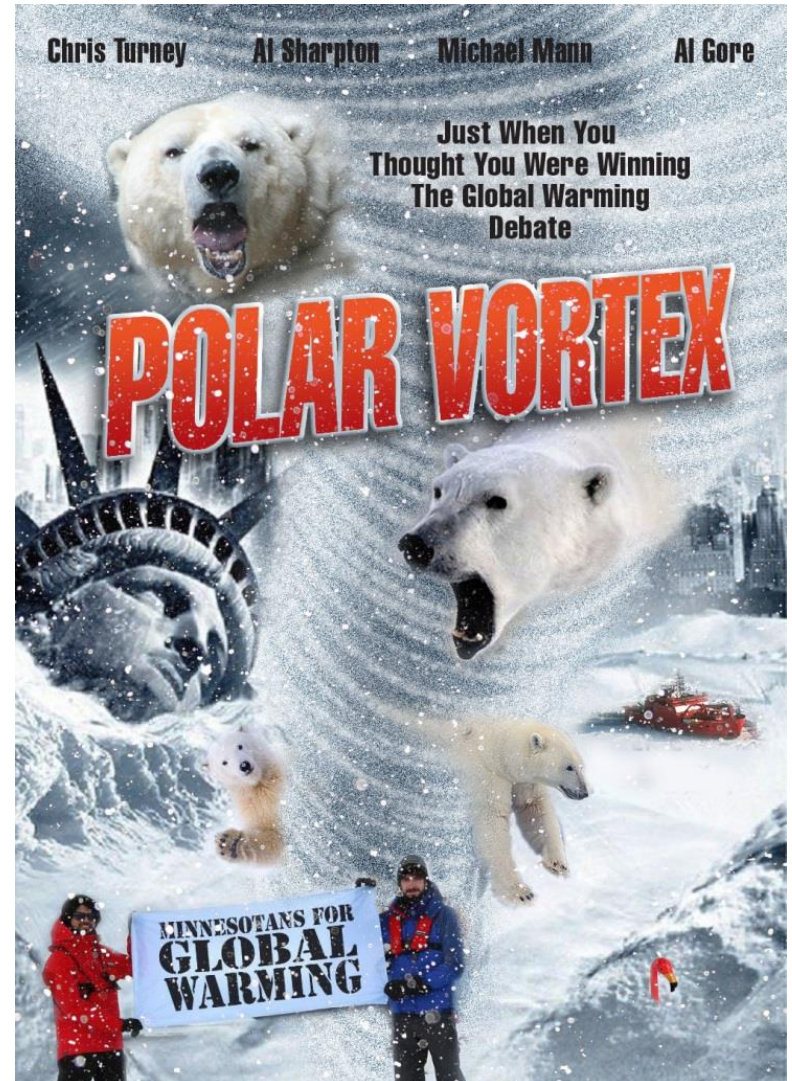
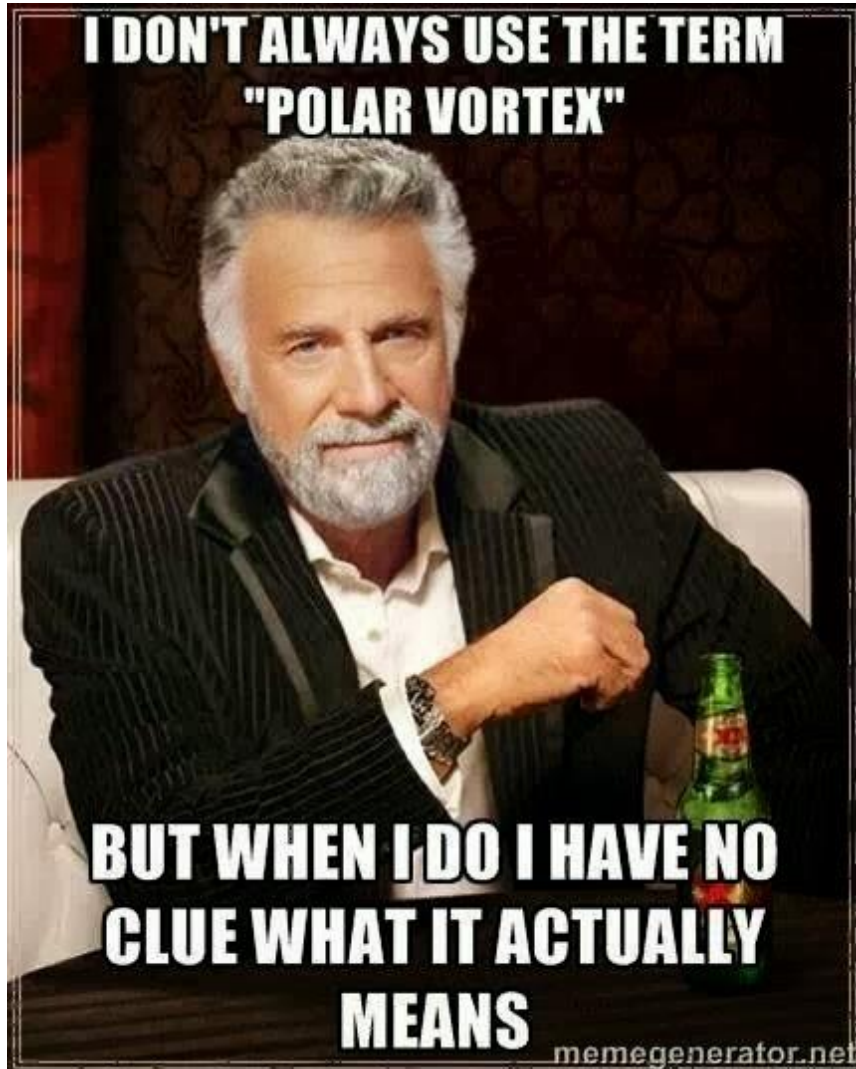


Polar Bears

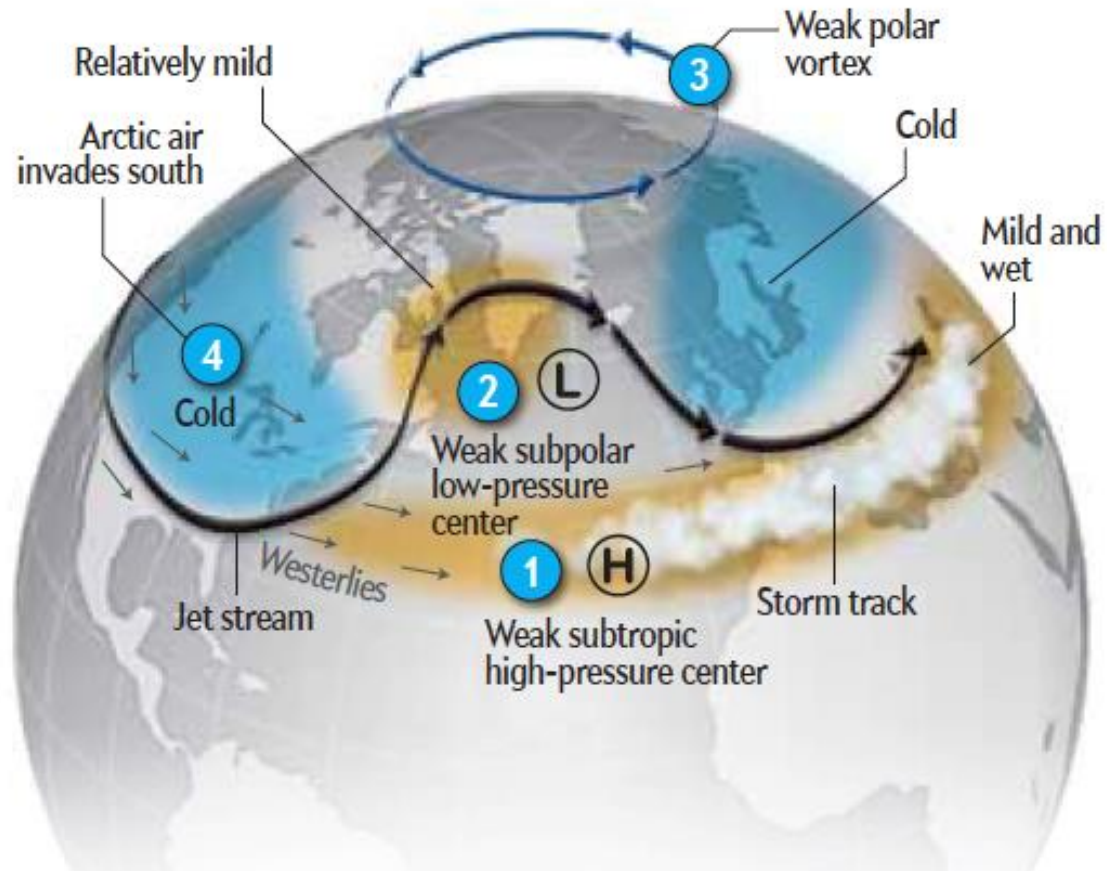
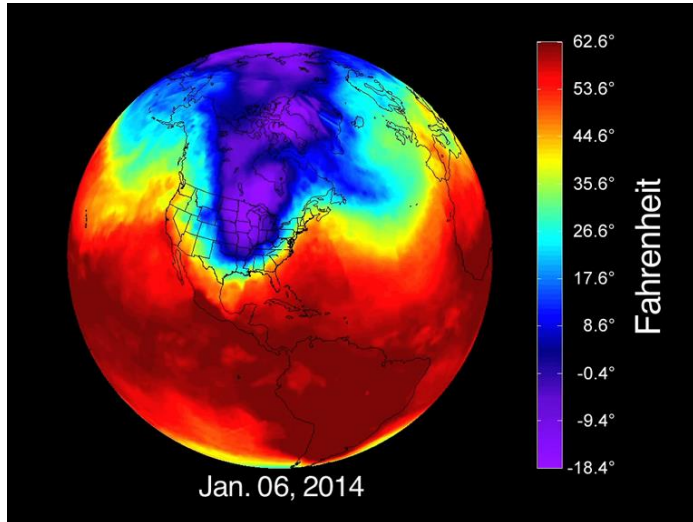
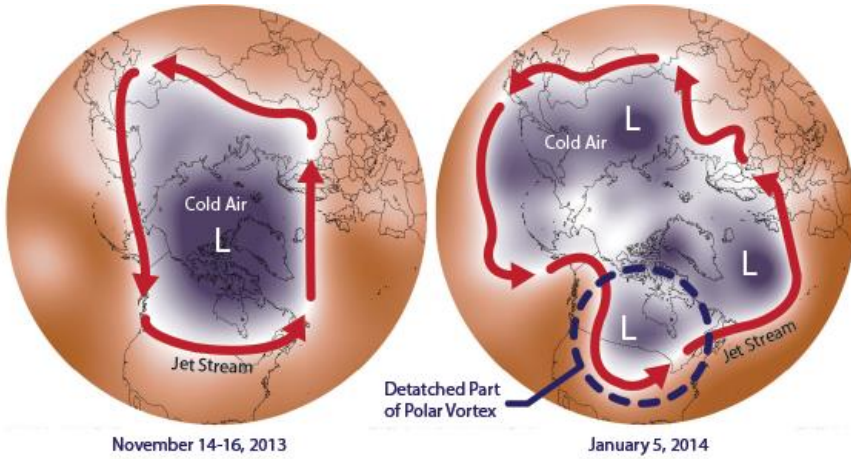
(all photos from Jay Roberts)



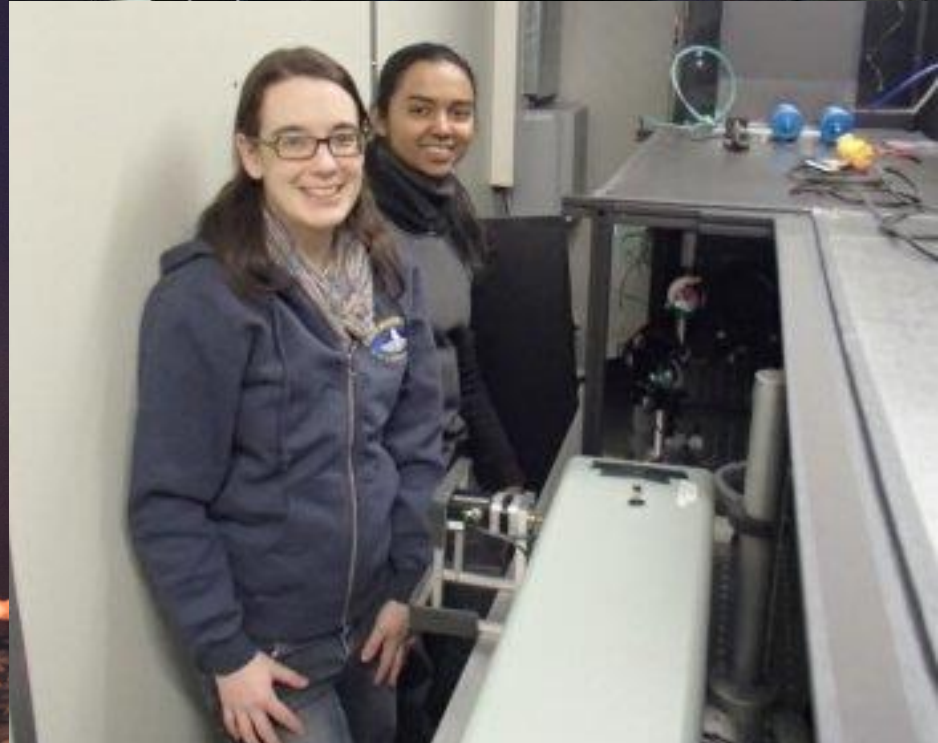
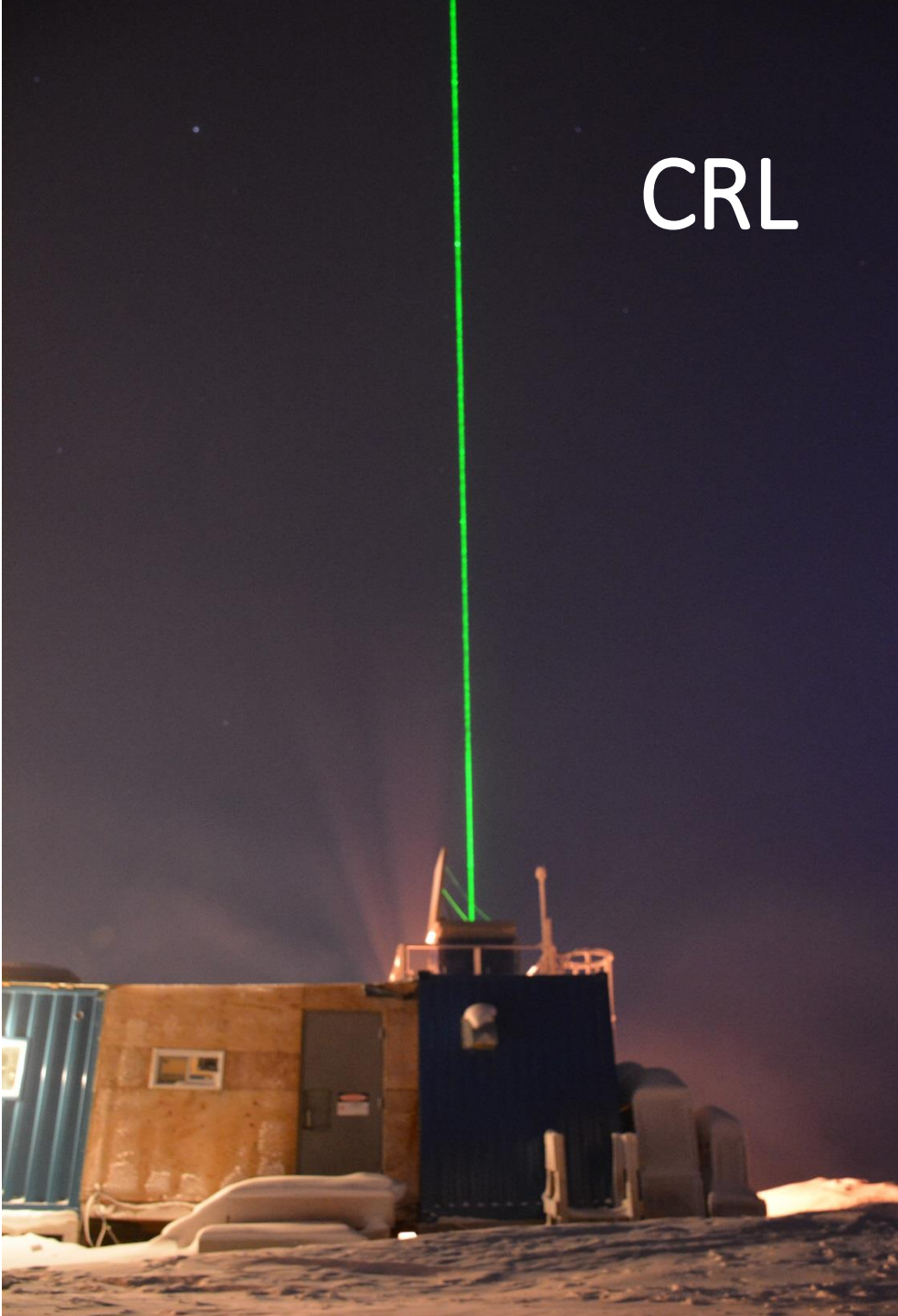
2014 Polar Vortex visits the south



Polar Vortex



CRL



Upper Paradise

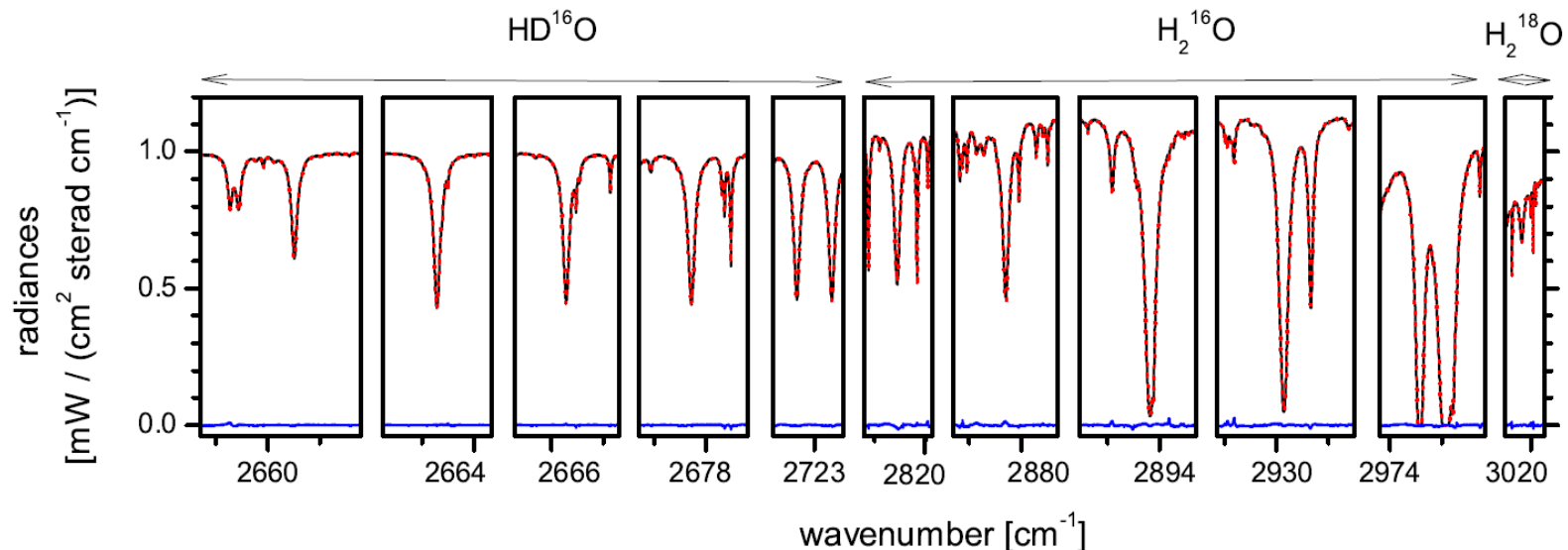


Line shapes

- Intrinsically, a spectral line should correspond to one unique energy value
- There are three primary causes of line broadening:
 - Natural broadening (molecule's energy uncertainty)
 - Has a negligible effect on the observed line shapes. .
 - Pressure broadening (caused by molecular collisions)
 - Dominant effect on the wings of spectra lines for molecules below ~20 km in altitude since it is a pressure-dependent phenomenon (Wallace & Hobbs, 2006).
 - Doppler broadening (caused by motion)
 - Dominant effect on line shapes for molecules above ~ 50 km, but is significant throughout the atmosphere, and is stronger in the center of a spectral line.
- This provides the ability to recover information about the distribution of gases in the atmosphere.
- Often, Pressure and Doppler broadening are combined to be described as a Voigt line shape.

MUSICA microwindows

- Cover a range of weak and strong absorption lines to gain sensitivity to the significant range of abundances.
- These microwindows have four interfering species: O_3 , CO_2 , N_2O , CH_4 .
 - Retrieval of CO_2 is used as a quality control test, as abundances are well-known.



MUSICA microwindows for ground based FTIR water vapour retrievals (Fig. 2, Schneider et al., 2012).